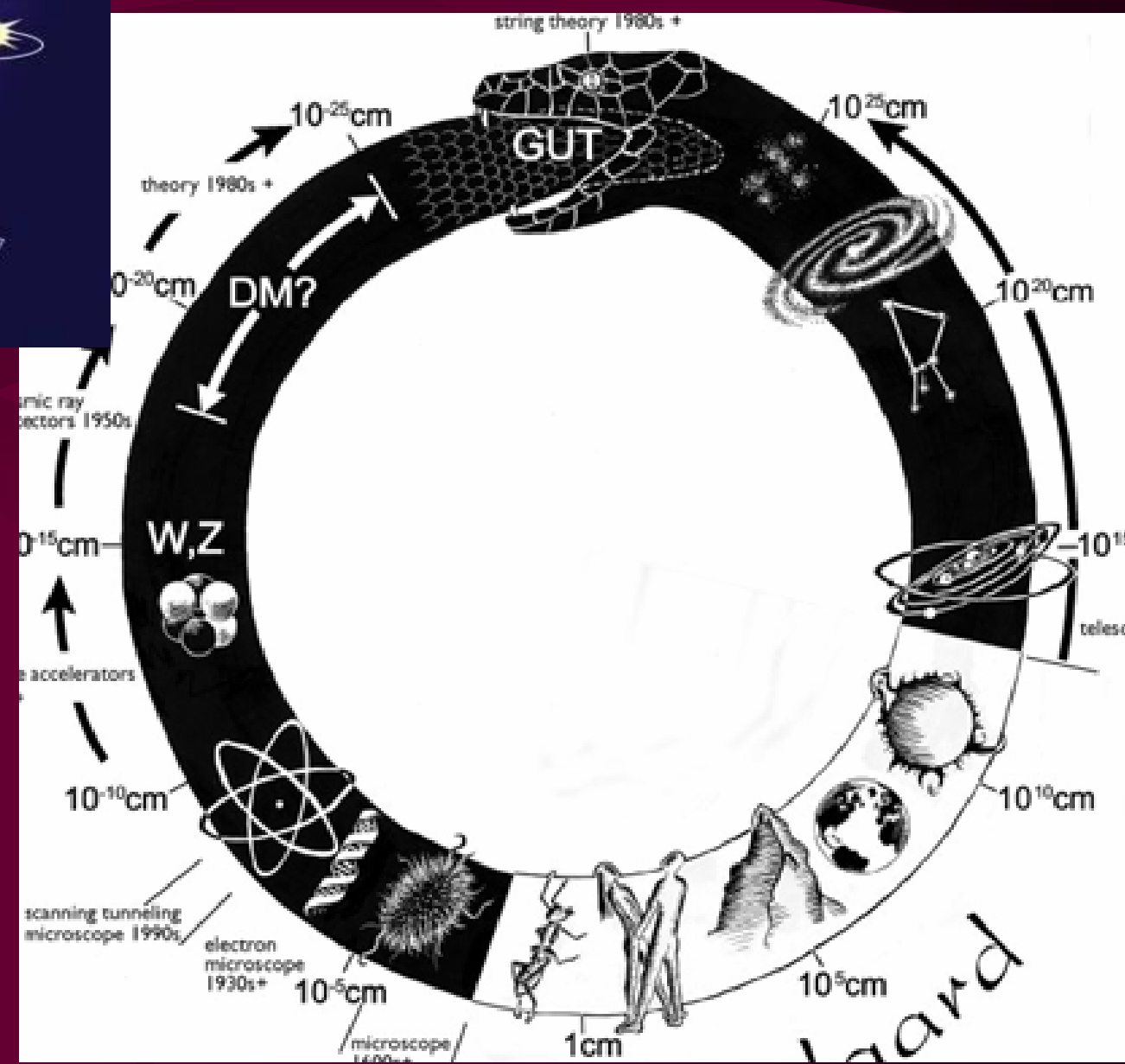
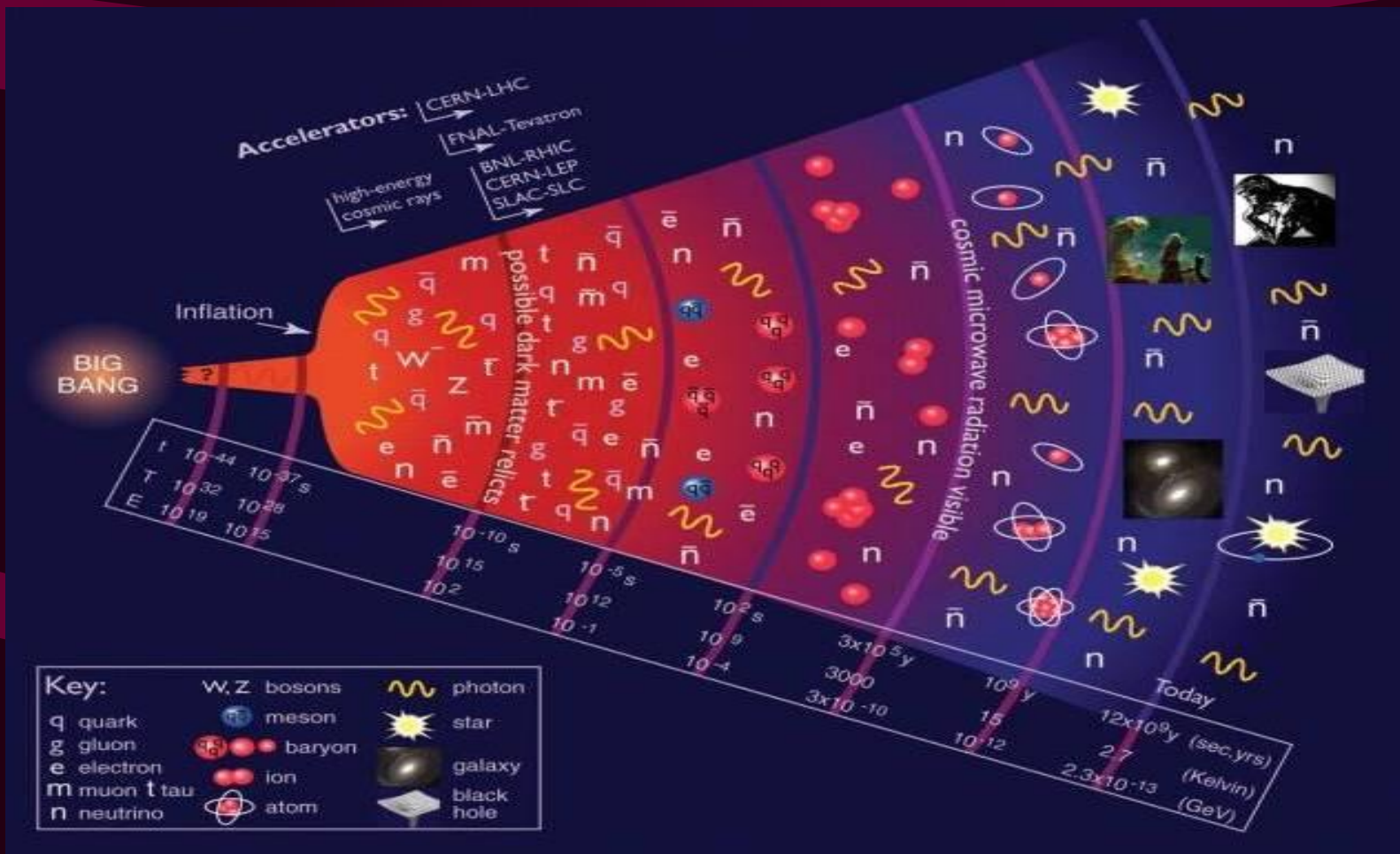


# Problematic *WMAP* Cosmology

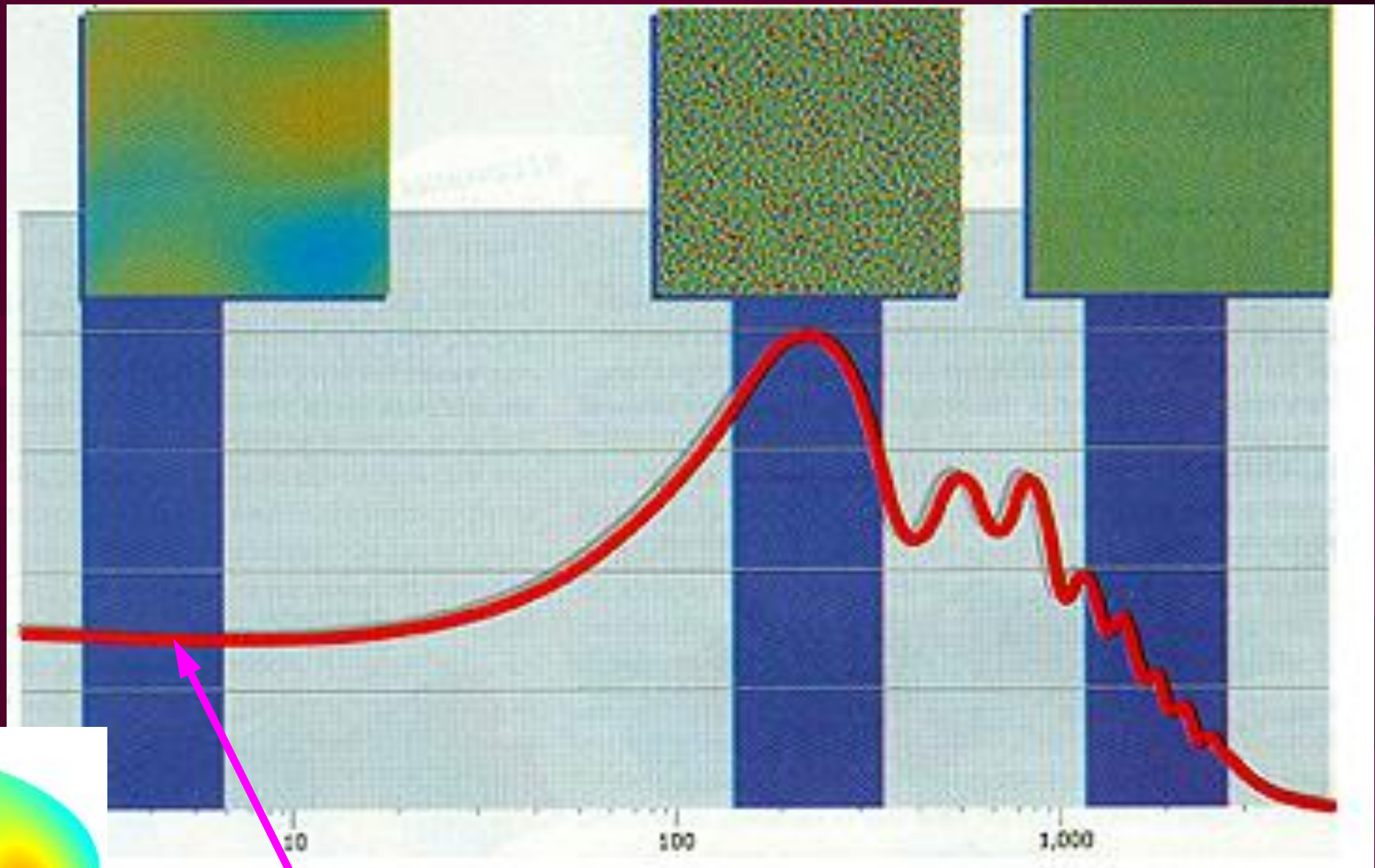
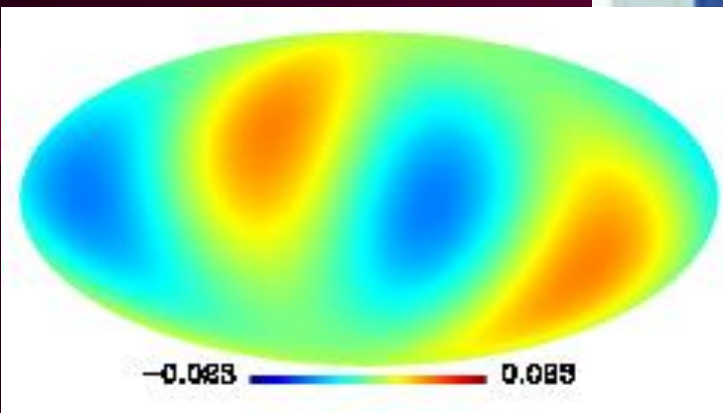
LI Ti-Pei

Department of Physics, Tsinghua University  
Institute of High Energy Physics, CAS

- I. Dipole effect on CMB map
- II. Artificial anisotropy in *WMAP* CMB map
- III. What about *Planck* ?
- IV. Impact on cosmology and physics



$$Q \approx 1000 \mu\text{K}^2$$

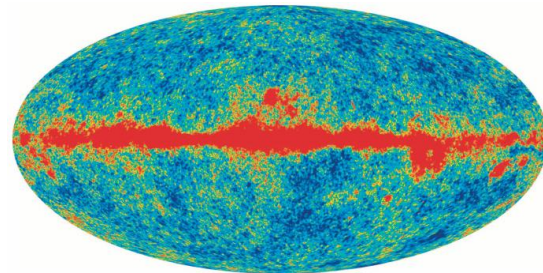


$\Lambda$ CDM

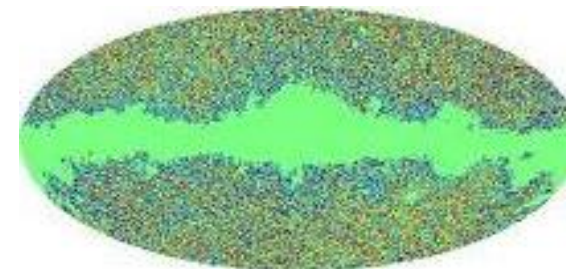
# I. Dipole Effect on CMB Map

## 1、Contaminations in CMB Experiments

### (1) Foreground



mask



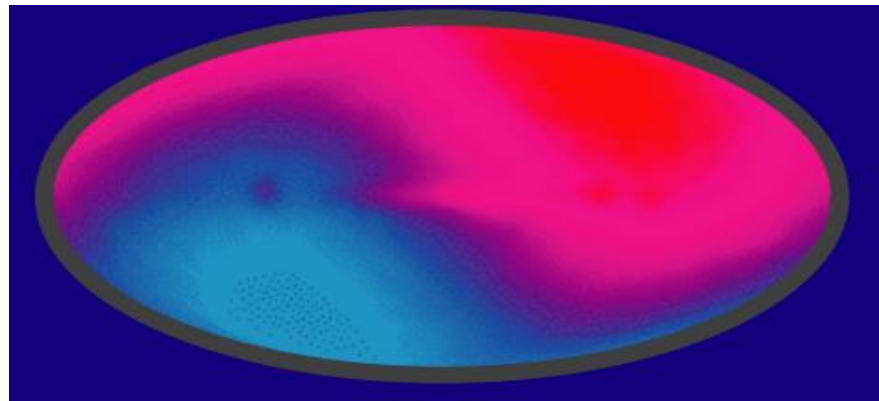
- **radio sources**
- **diffuse foreground**

<u>process</u>	<u>template</u>	<u>template fits</u>
dust emission	$T_d$	$T_{clean} = T_{obs} - (c_d T_d + c_H T_H + c_s T_s)$ $\text{Var}(T_{clean}) = \min \longrightarrow c_d \ c_H \ c_s$ “aesthetic criteria”
free-free emission	$T_H$	
synchrotron emission	$T_s$	

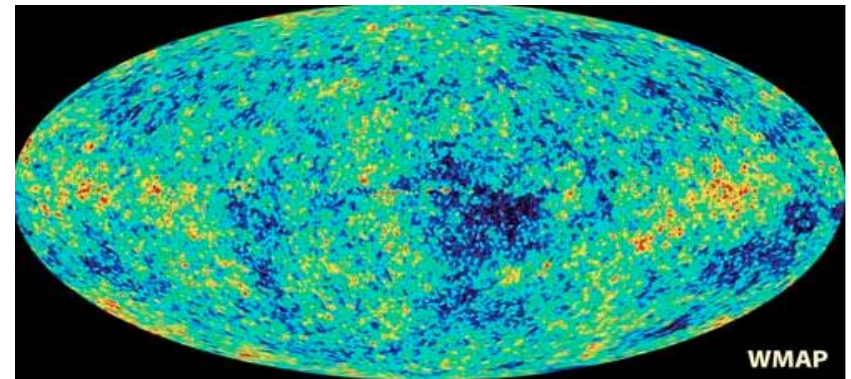
## (2)、Dipole contamination

- Dipole's amplitude

$$T_{dipole}(t) = \frac{T_0}{c} \mathbf{v}(t) \cdot \mathbf{n}(t)$$



$$\Delta T_d \sim 3 \text{ mK}$$



$$\Delta T_{CMB} \sim 50 \mu\text{K}$$

### (3) Error in calculated dipole $T_d = \frac{T_0}{c} \mathbf{v} \cdot \mathbf{n}$

- CMB dipole direction

$$\Delta T_d = \frac{T_0}{c} \Delta \mathbf{v} \cdot \mathbf{n}$$

$$\Delta \mathbf{v} \sim 10' \Rightarrow \Delta T_d \sim 30 \mu\text{K} \quad (\Delta T_{CMB} \sim 50 \mu\text{K})$$

- Sidelobe contamination

$$T_s(p(t)) = \sum_{p'_s} G'(p'_s) T_d(p_s) / N \quad T_{CMB}(p) = T(p) - T_s(p)$$

$$\Delta T_s(p(t)) = \frac{T_0}{c} \mathbf{v}'(t) \cdot \sum_{p'_s} \frac{\Delta G'(p'_s)}{N} \mathbf{n}'(p'_s) = \frac{T_0}{c} \mathbf{v}'(t) \cdot \Delta \mathbf{n}'_s$$

$$\Delta \mathbf{n}'_s \sim 10'$$

## (4) Remove dipole-error-induced deviation

- Produce template maps

$\Delta T_x$  : take  $\Delta n = \Delta n_x = 1'$ , along the scan path  $\{p(t)\}$

calculate temperature deviation  $\Delta T_x(p) = -\frac{T_0}{c} \mathbf{v}(t) \cdot \Delta \mathbf{n}_x$

$\Delta T_y$  :

$\Delta T_z$  :

- Produce cleaned map

$$T_{clean} = T_{obs} - (c_x \Delta T_x + c_y \Delta T_y + c_z \Delta T_z)$$

$$\text{Var}(T_{clean}) = \min \longrightarrow c_x \quad c_y \quad c_z$$

“aesthetic criteria”

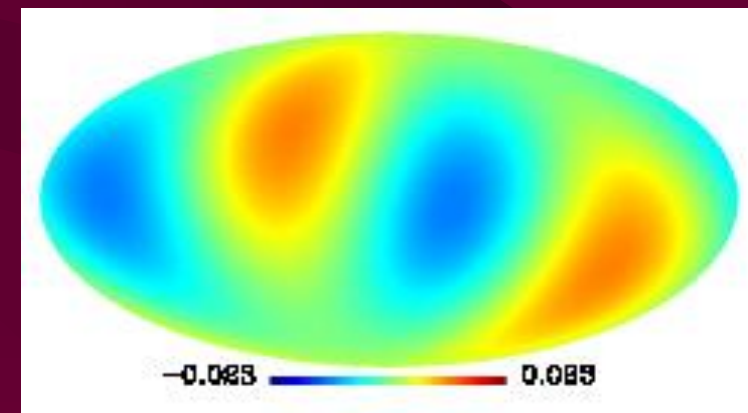
After correcting the timing error

→  $Q_{Clean} \approx (-3.2 \pm 3.5) \mu\text{K}^2$

WMAP raw data + (our pipeline + 7' offset)

→ WMAP release

Axis of Evil



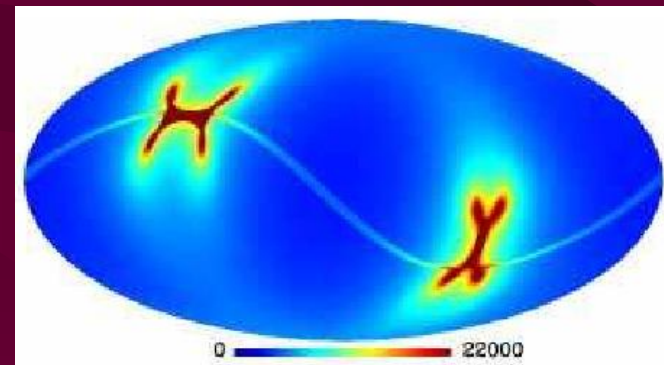
$Q_{WMAP} \approx 113 \mu\text{K}^2$



### III. What about *Planck* ?

Can *Planck* abstain from the artificial anisotropy induced by **dipole error + scan** ?

- timing/pointing
- **dipole direction**
- **sidelobe contamination**



## OBSERVATIONAL SCAN-INDUCED ARTIFICIAL COSMIC MICROWAVE BACKGROUND ANISOTROPY

HAO LIU<sup>1</sup> AND TI-PEI LI<sup>1,2,3</sup>

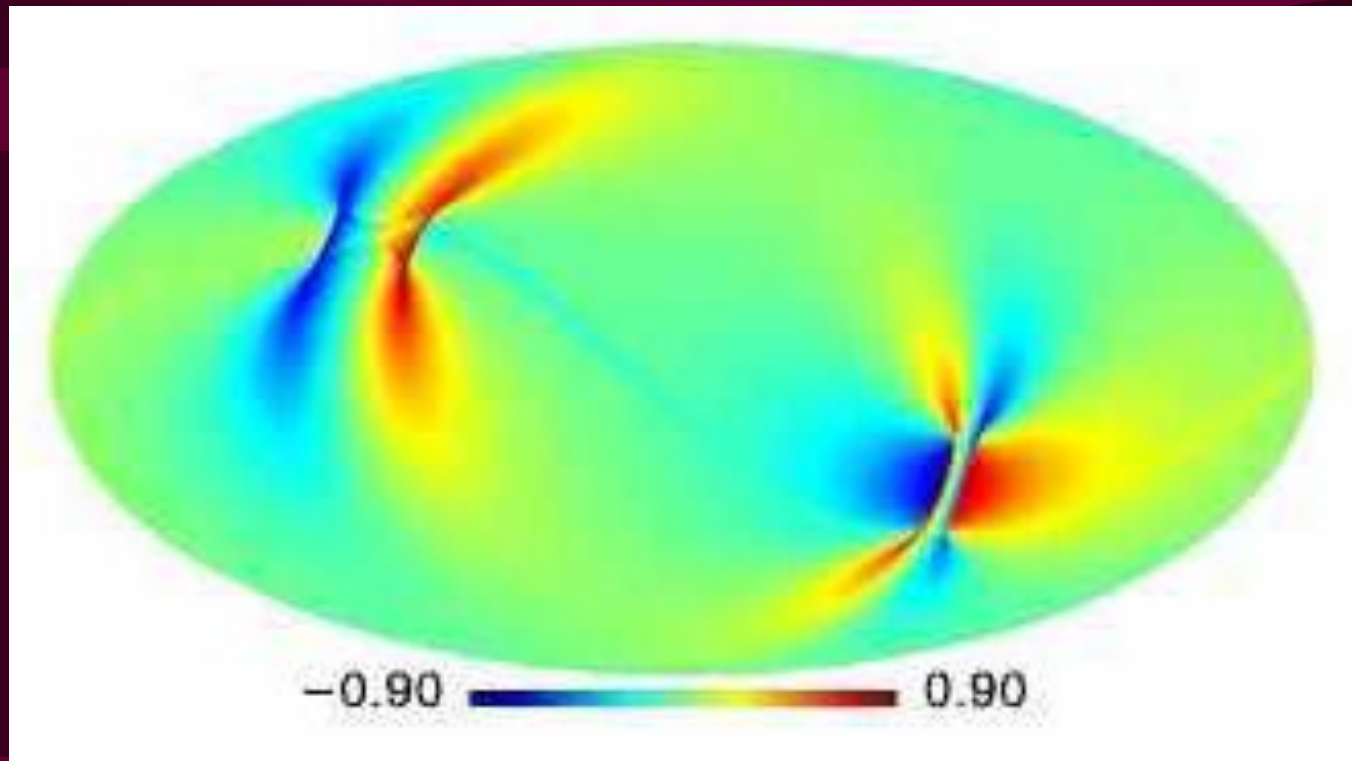
<sup>1</sup> Key Laboratory of Particle Astrophysics, Institute of High Energy Physics, Chinese Academy of Sciences, Beijing, China; [liuhao@ihep.ac.cn](mailto:liuhao@ihep.ac.cn)

<sup>2</sup> Department of Physics and Center for Astrophysics, Tsinghua University, Beijing, China; [ltp@tsinghua.edu.cn](mailto:ltp@tsinghua.edu.cn)

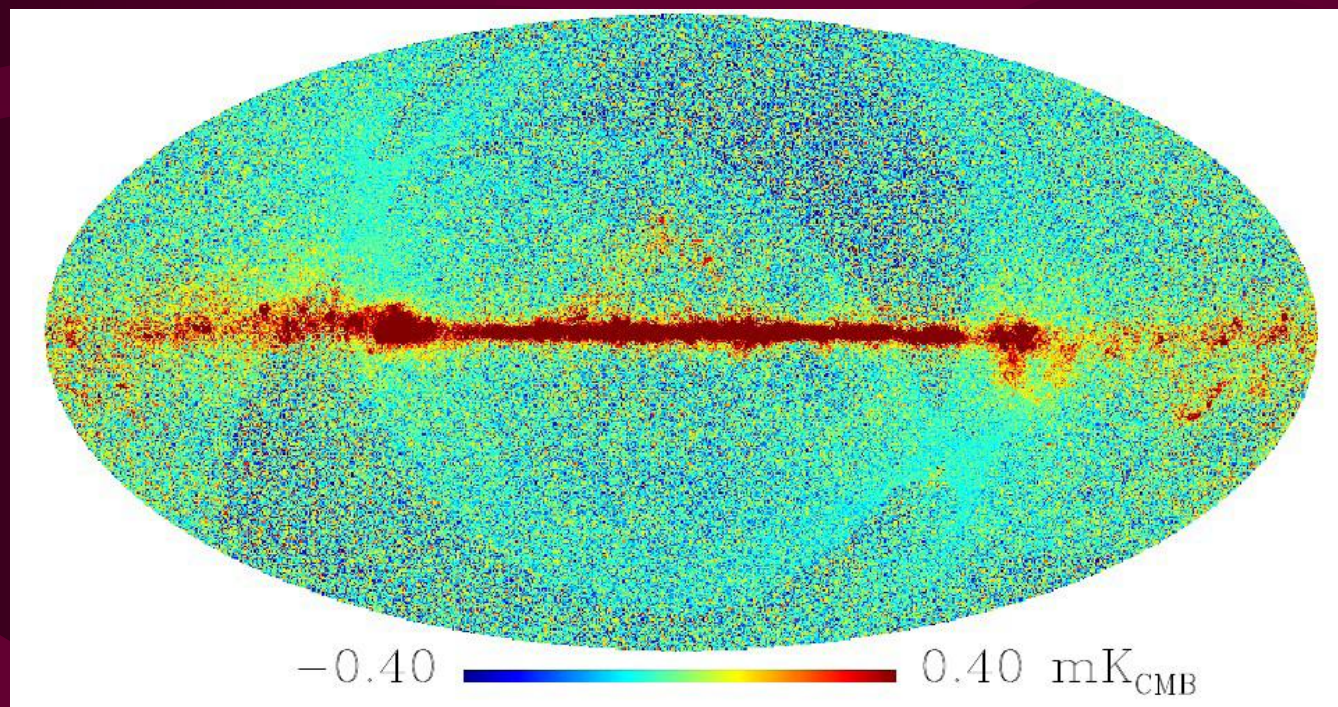
<sup>3</sup> Department of Engineering Physics and Center for Astrophysics, Tsinghua University, Beijing, China

Received 2011 January 10; accepted 2011 March 15; published 2011 April 26

“The scan-induced anisotropy is a common problem for all sweep missions and, like the foreground emissions, has to be removed from observed maps. Without doing so, CMB maps from *COBE*, *WMAP*, and *Planck* as well, are not reliable for studying the CMB anisotropy.”



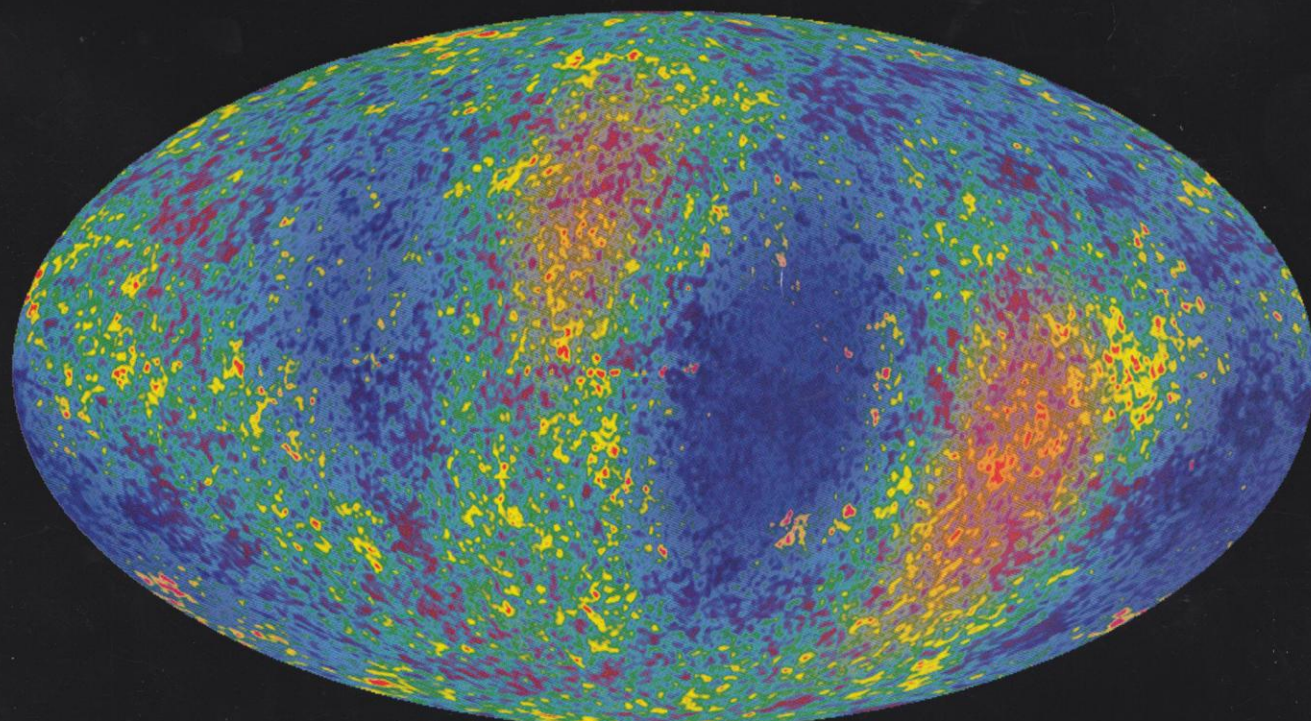
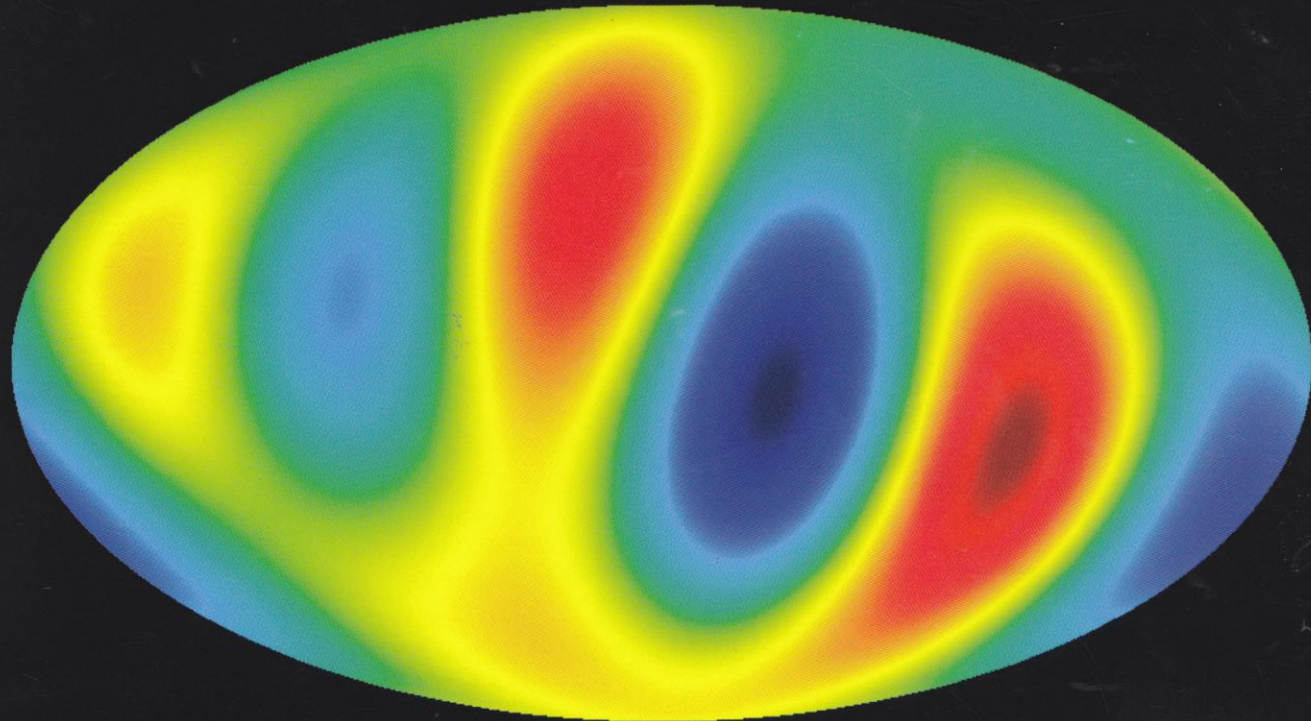
*Liu & Li 2011 ApJ*



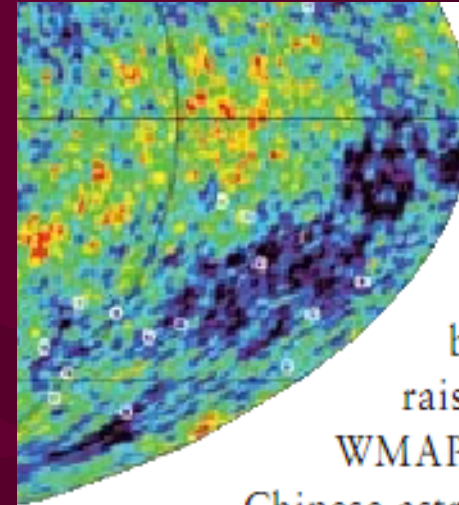
*Planck 2011*

# IV. Impact on Cosmology and Physics

## WMAP cosmology questioned

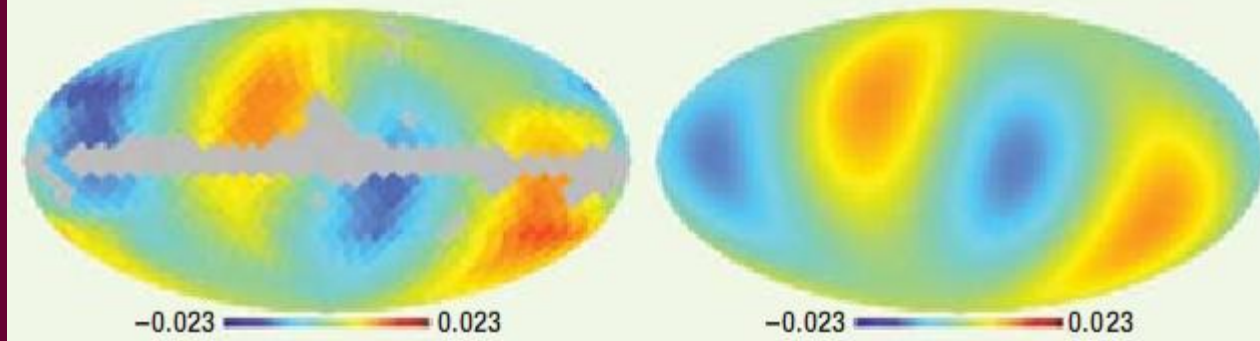


## Is everything we know about the universe wrong?

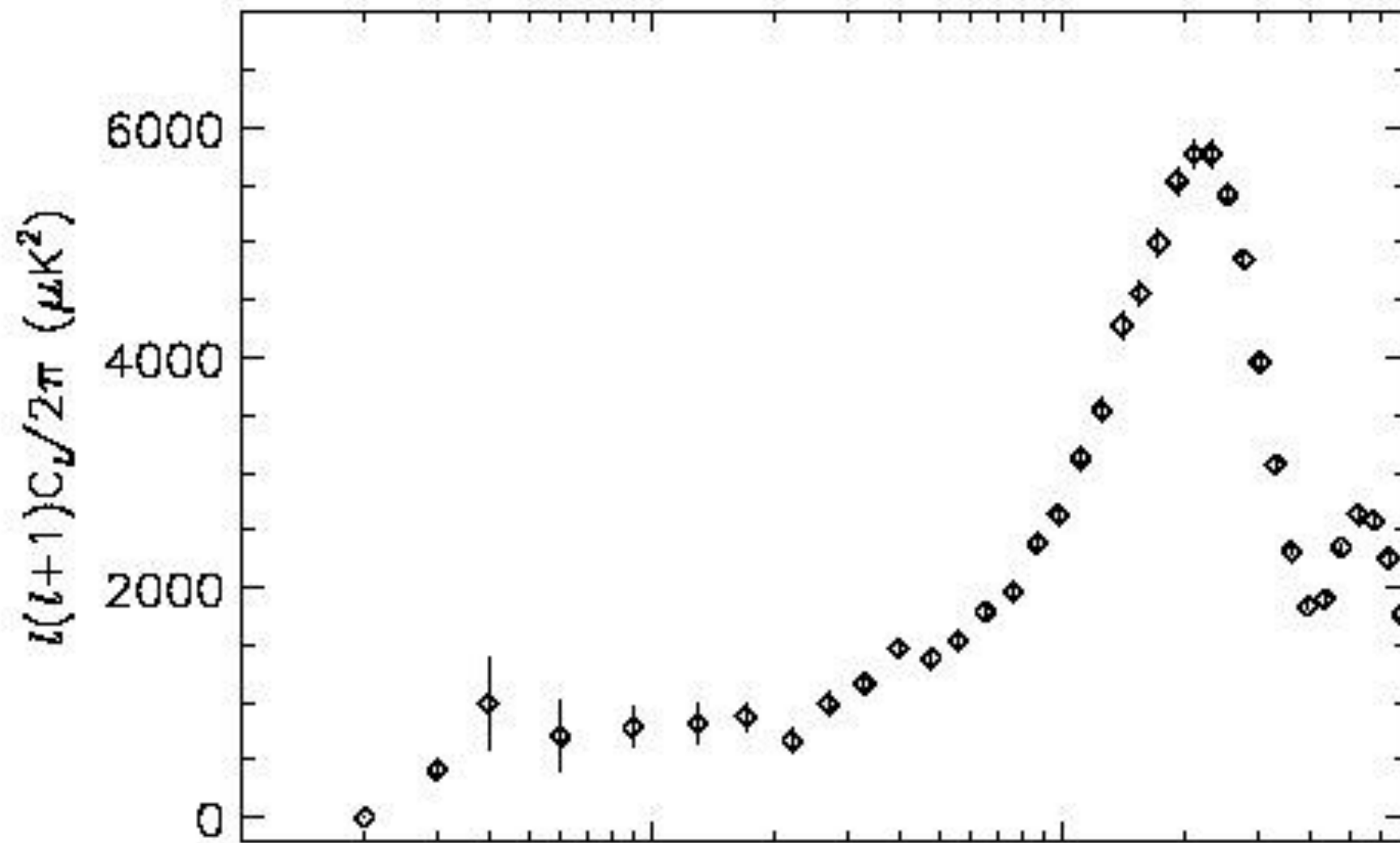


### Axis of evil and 25.6 ms

Since our WMAP beam paper was submitted there have been other criticisms raised about the official WMAP data reduction. Two Chinese astronomers, Liu and Li (2010), have suggested that confusion over a tiny 25.6 ms offset between recorded timings of where WMAP was pointing and when the CMB temperature was measured may be enough to explain several anomalies found at very large scales in the official WMAP maps. Various authors have remarked that the largest scale components (quadrupole etc) align with the ecliptic. These anomalies are usually termed



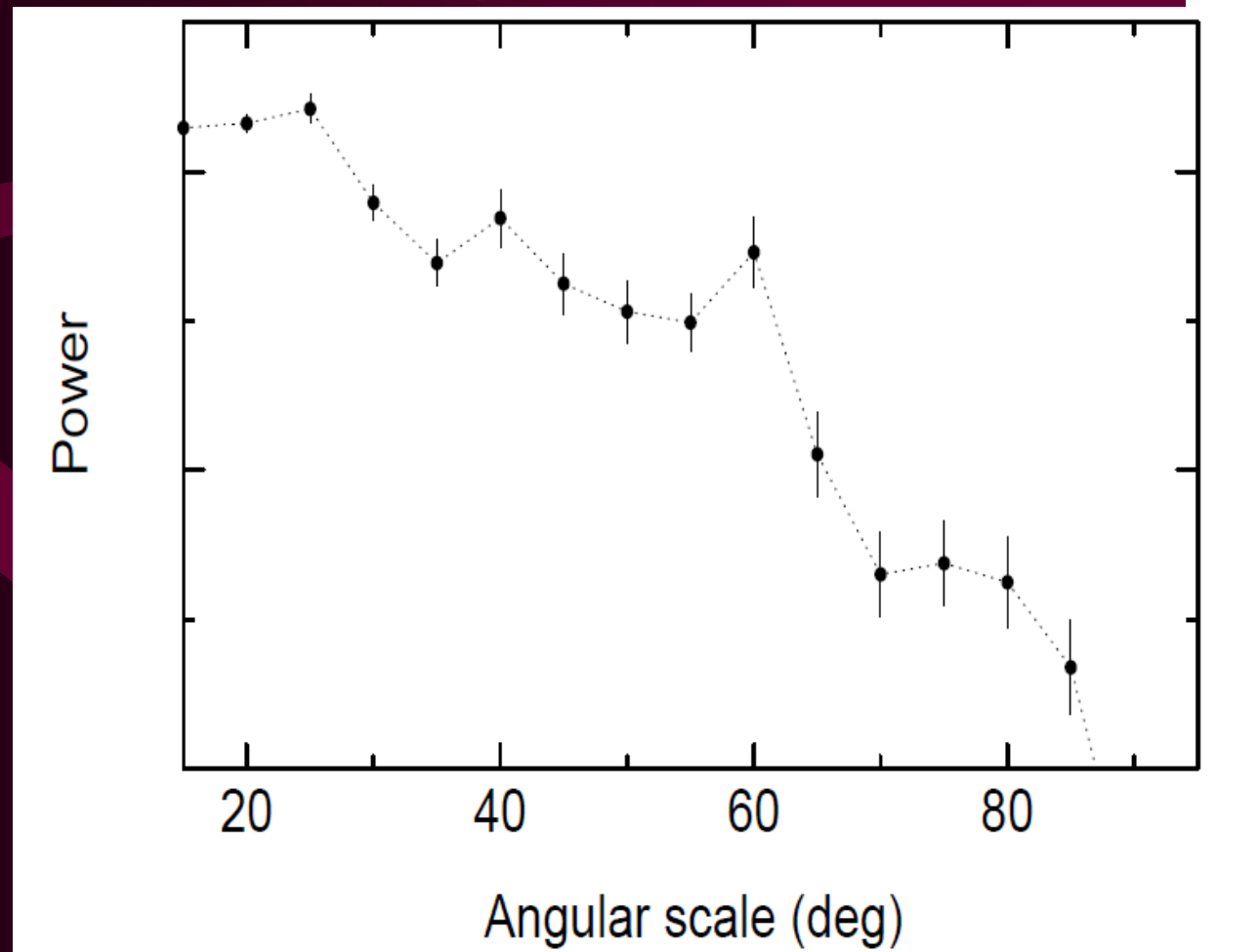
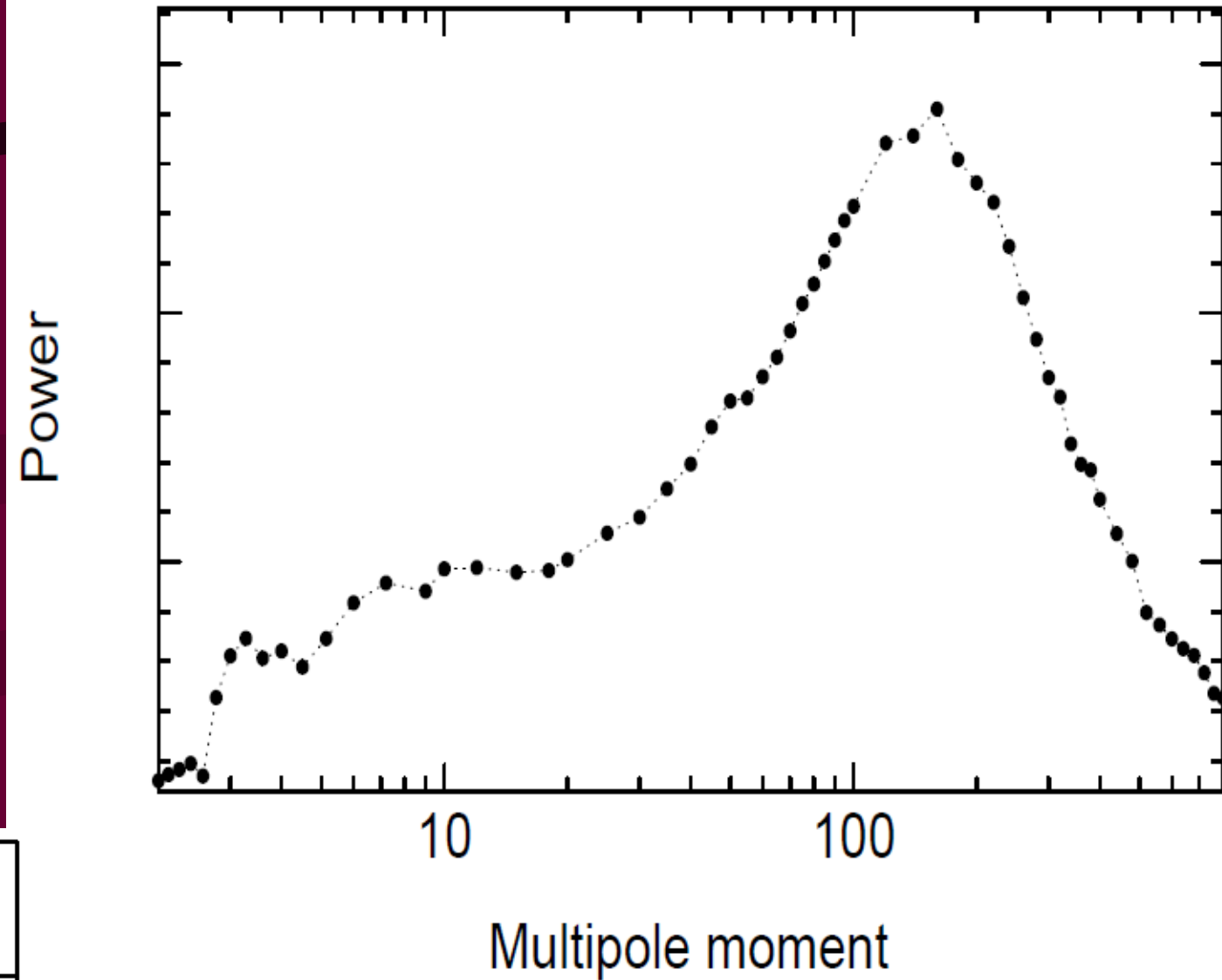
3 (Left): The temperature map produced from the Liu and Li (2010) simulation comprising only a dipole and map-making software including the 25.6ms timing offset with no inclusion of primordial fluctuations. (Right): The official WMAP CMB large-scale (quadrupole) "map". Both panels are in galactic coordinates and in units of mK. There is a remarkable similarity between the large-scale (quadrupole) features in both maps even though the simulated map knows nothing about the real universe! (Liu and Li 2010)



$$Q_{Clean} \approx (-3.2 \pm 3.5) \mu\text{K}^2$$

→ Cold and stable very early universe  
no quantum fluctuation?

# Large scale CMB — a window to the early universe



- Accelerating Universe

⇒  $G + \Lambda$  in the present universe

- Zero Quadrupole

⇒  $\Lambda + G$  in the very early universe



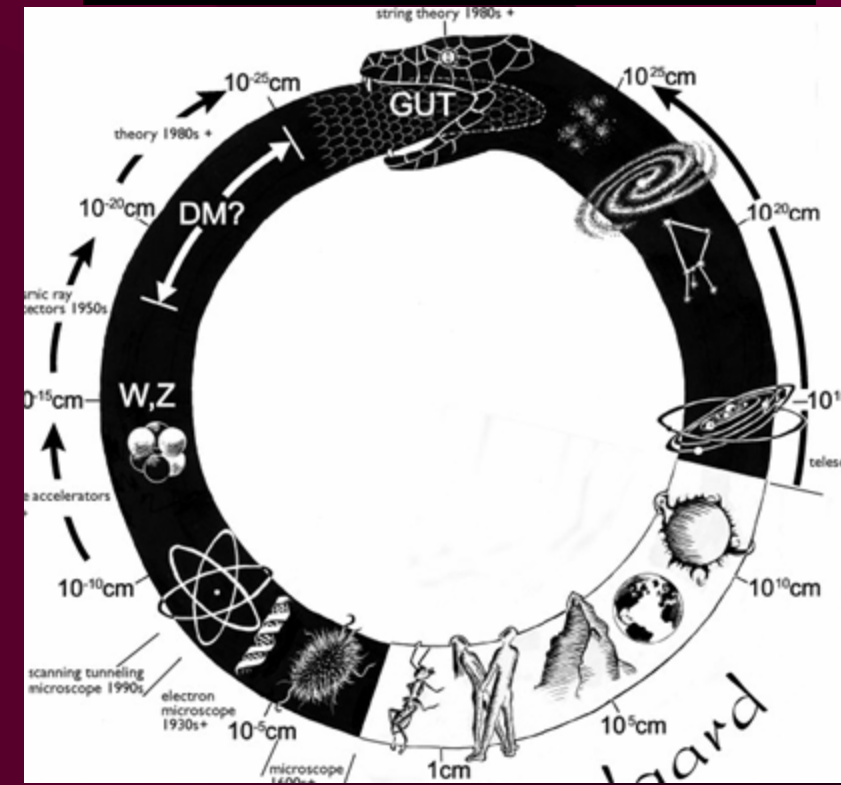
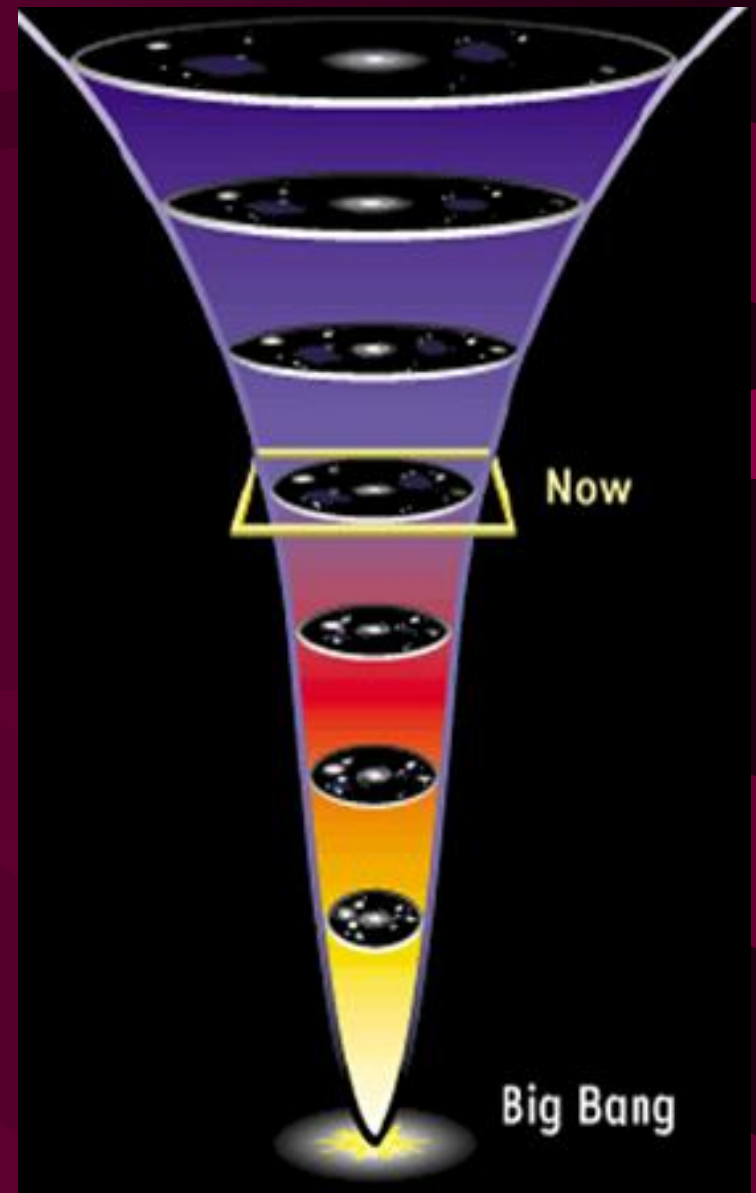
# 3、新物理

大统一？

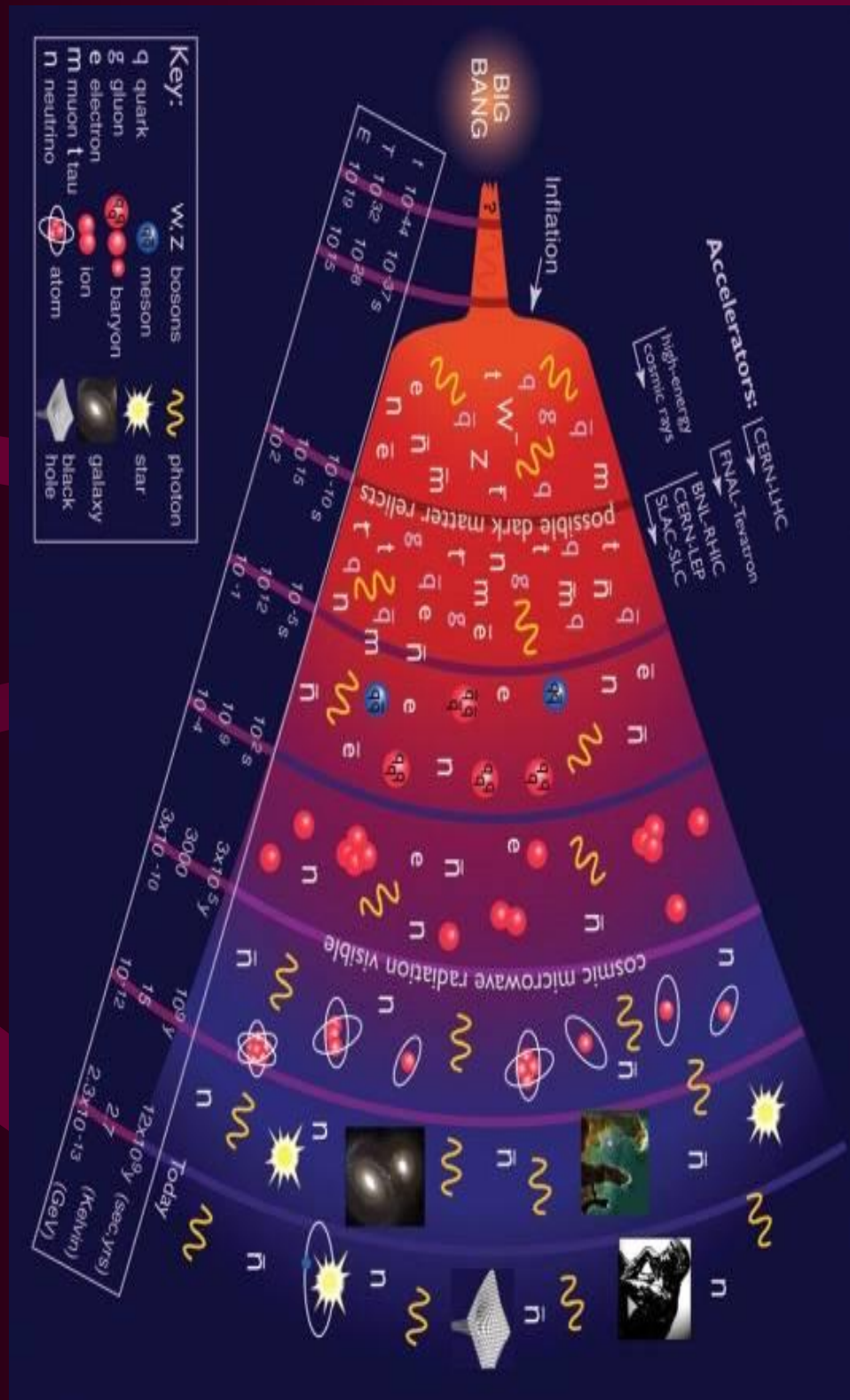
泰勒斯：万物皆由水构成

赫拉克利特：

从一切产生一，从一产生一切

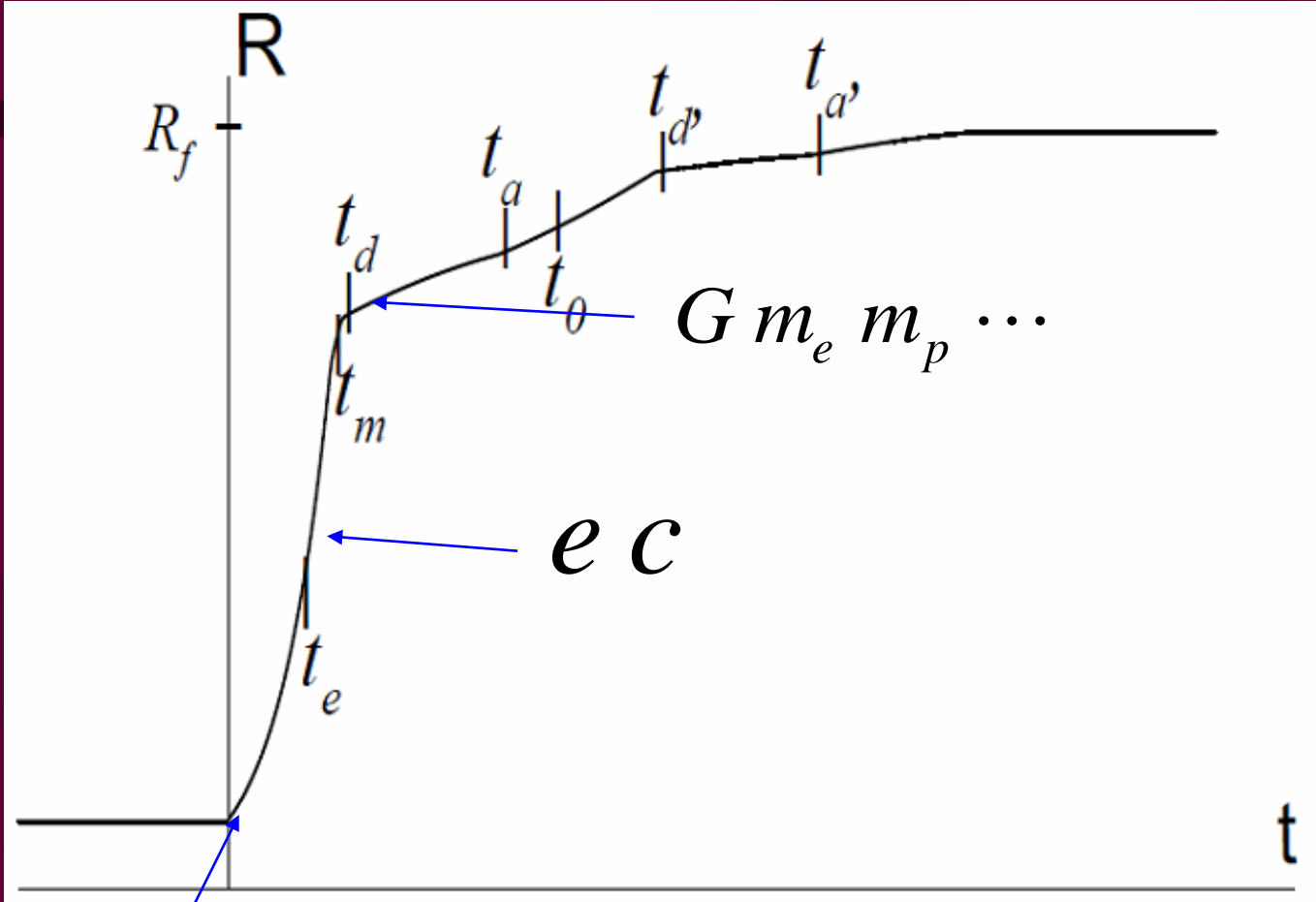


# 宇宙的热历史

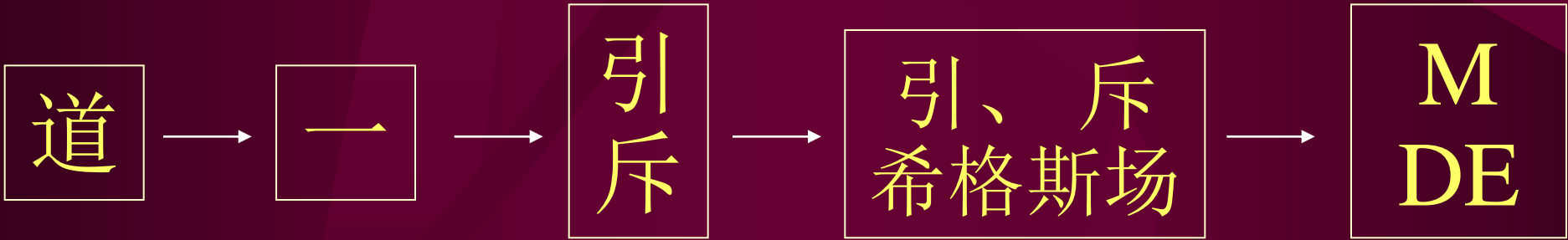


温度 (K)	能量 (eV)	时间 (秒)	时代	物理过程
$10^{32}$	$10^{28}$	$10^{-44}$	Planck	
$10^{28}$	$10^{24}$	$10^{-36}$	大统一	
		$10^{-35,-33}$	暴胀	暴胀
$10^{15}$	$10^{11}$	$10^{-10}$	电弱统一	
$10^{13}$	$10^9$	$10^{-6}$	强子	夸克禁闭
$10^{11}$	$10^7$	$10^{-2}$	轻子	
$10^{10}$	$10^6$	1	弱力出现	中微子脱耦
$5 \times 10^9$	$5 \times 10^5$	5	电子对湮灭	电子对湮灭
$10^9$	$10^5$	$10^2$ (3分)	核合成	轻核素生成
$3 \times 10^3$	0.3	$10^{13}$ (30万年)	复合	物质 - 辐射退耦 宇宙背景辐射
		$10^{16}$ (4亿年)	第一代恒星生成	再电离
			星系; 星系团	大尺度结构形成 开始加速膨胀
2.7	$3 \times 10^{-4}$	$4 \times 10^{17}$ (137亿年)	现代	

道生一  
 一生二  
 二生三  
 三生万物  
 万物负阴而抱阳  
 冲气以为和



$R_i \Lambda_a \Lambda_r$  : 随机



“统一” 新物理  
 A R (h)

*Thank you*



**Key Lab. of Particle Astrophysics IHEP/CAS**



**THCA / Tsinghua University**