

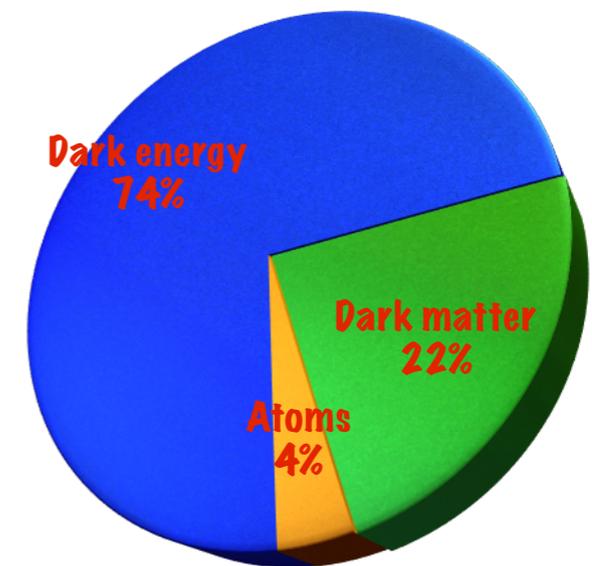
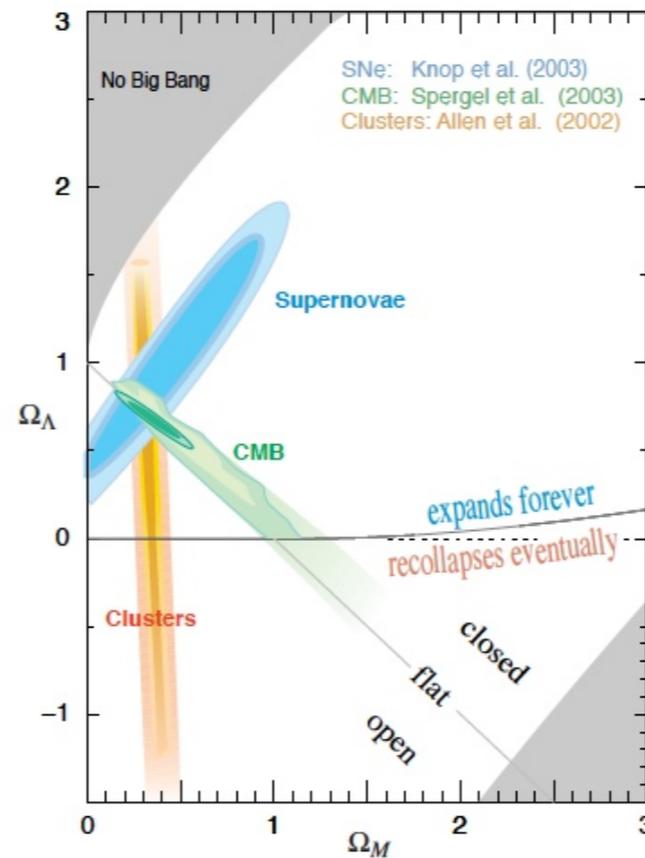
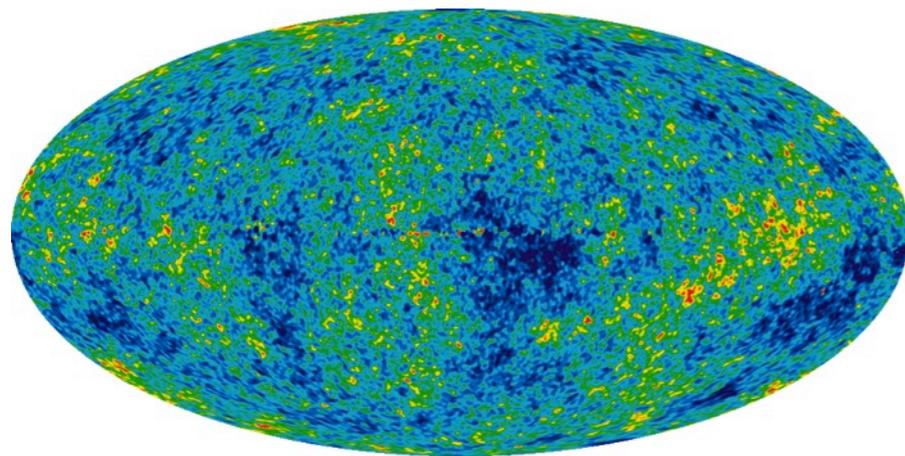
Isospin Violating Dark Matter and Neutrinos From the Sun

陈绍龙

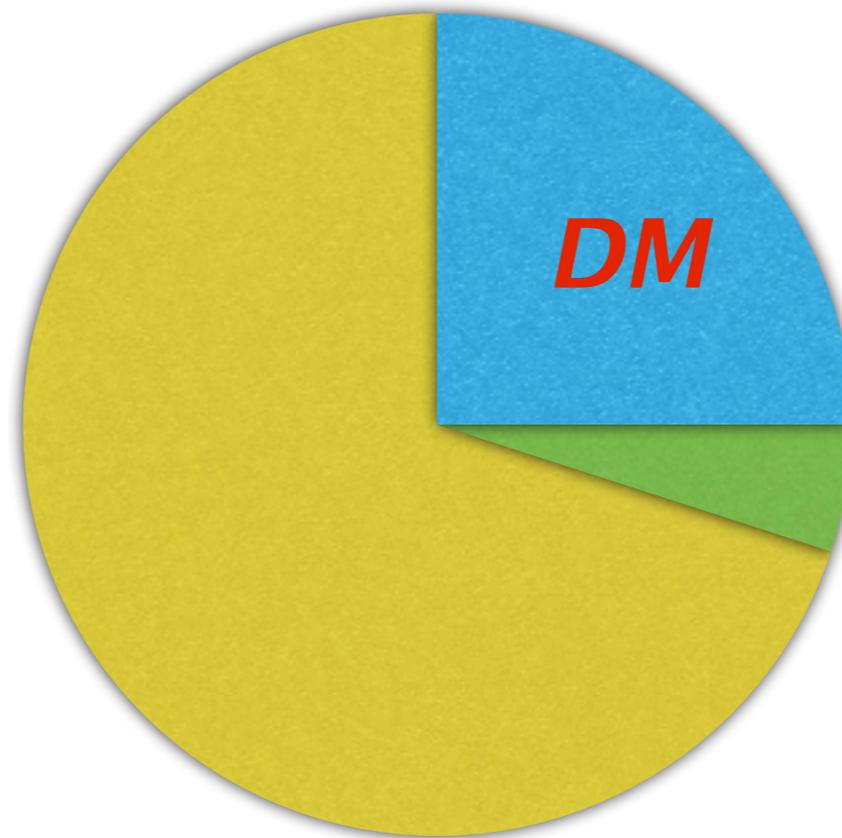
华中师范大学

两岸粒子物理与宇宙学研讨会，重庆，2012年5月

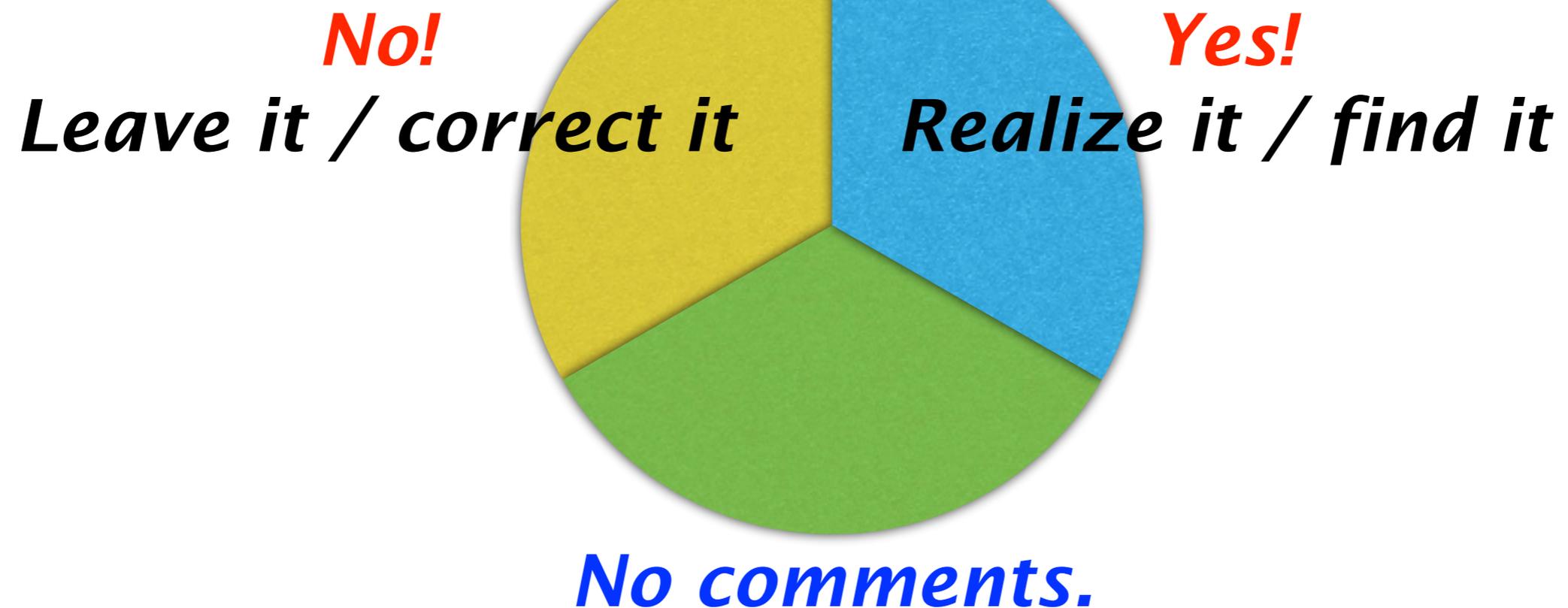
Evidence for DM (?)



Question: Do you believe in Dark Matter?



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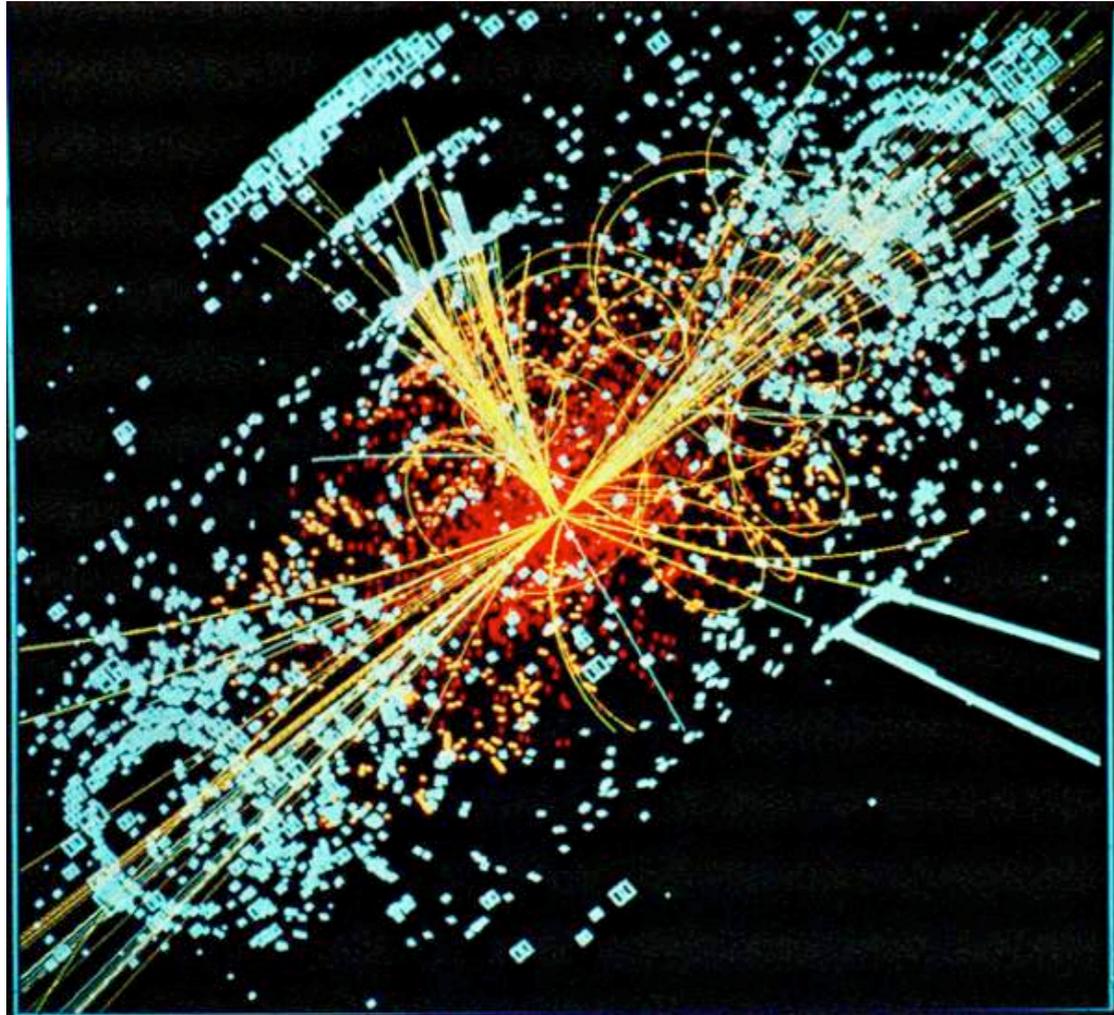


- **Basic properties:** Neutral, Massive, Non-Baryonic, Weakly interacting, Stable/long-lived.
- **Popular Candidates:** Weakly interacting Massive Particles (WIMPs).
- **Predicted (Embedded) in many models Beyond the SM** addressing the hierarchy problem. (TeV/weak scale DM)
- **WIMP miracle:**

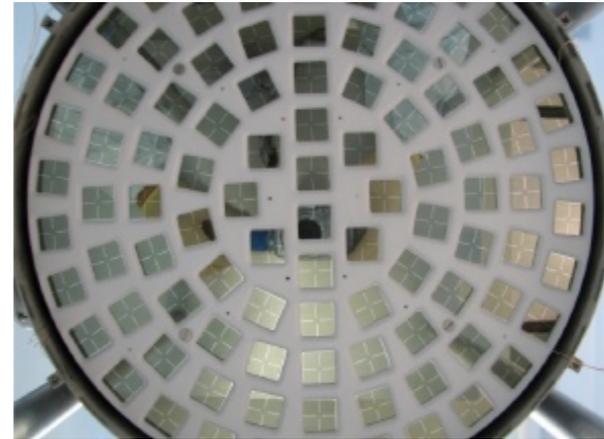
$$\Omega_\chi h^2 \simeq \frac{3 \times 10^{-27} \text{cm}^3/\text{sec}}{\langle \sigma v \rangle} \sim 0.11 \quad \Leftrightarrow \quad \sigma \sim \frac{g^4}{M_{\text{weak}}^2} \sim 1 \text{pb}$$

- **Direct detection era:**

$$\sigma_{DM_N} \lesssim (\sim) 10^{-44} \text{cm}^2$$



对撞机产生



直接探测



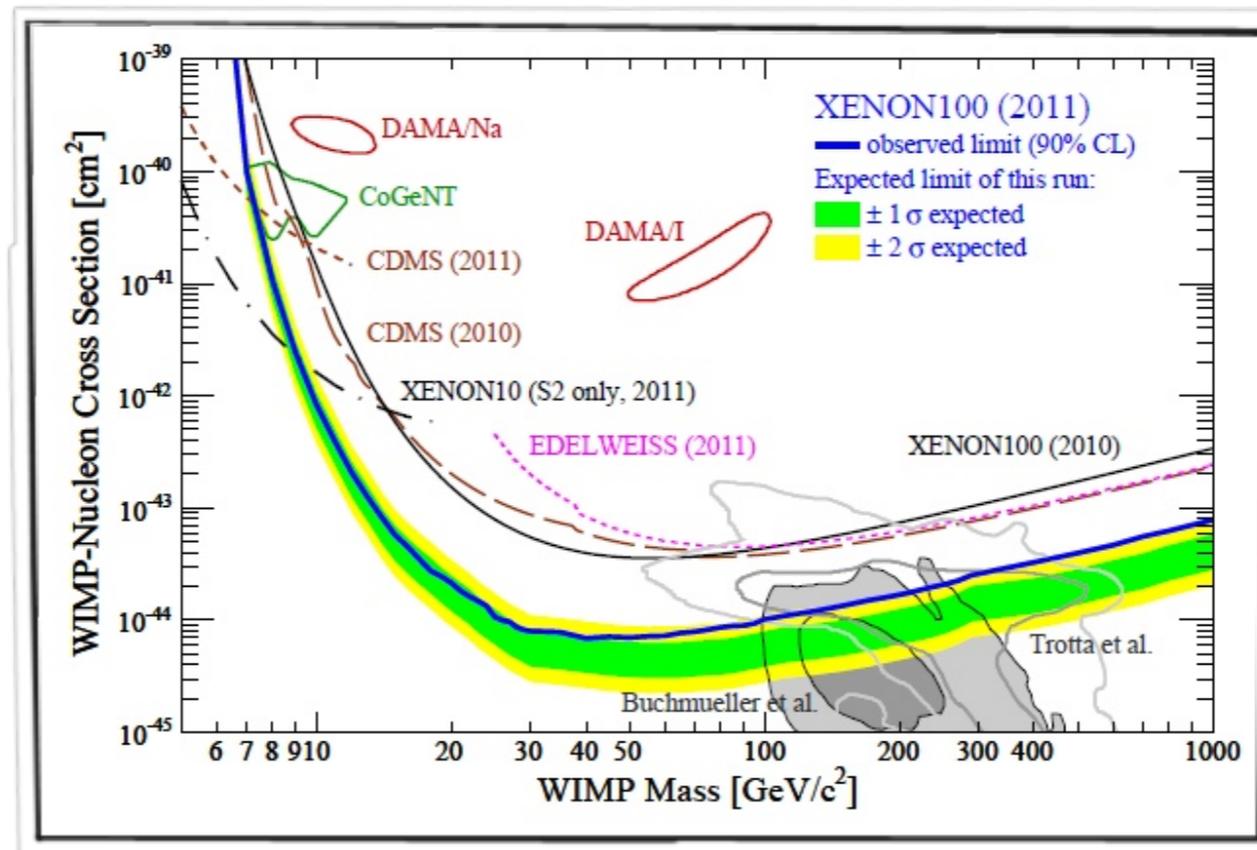
间接探测

- Indirect detection (ATIC, FERMI, HESS, PAMELA, ...)
Positron and electron excess

⇒ **TeV/weak scale Dark matter explanation**

- Direct detection:

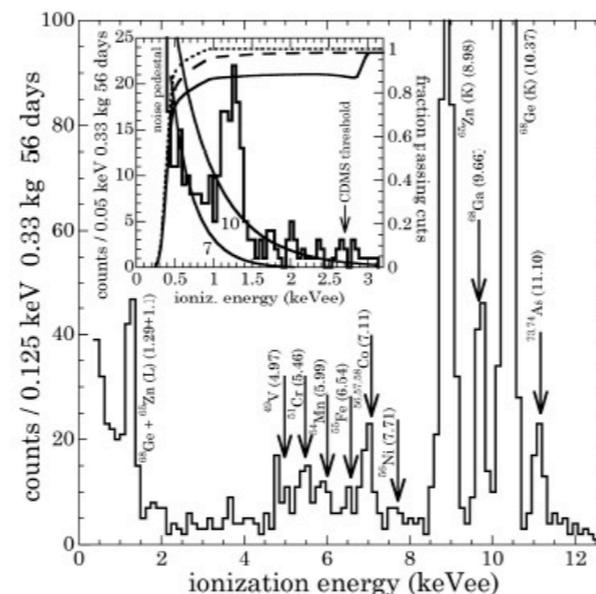
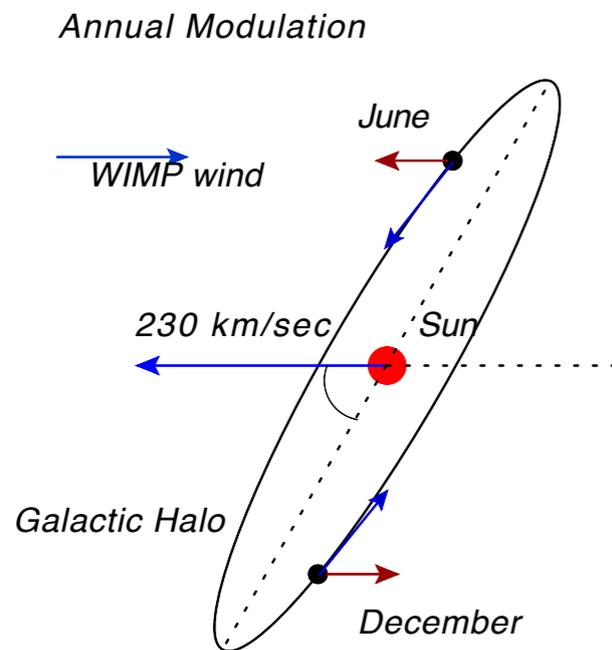
Null Experiments (CDMS, XENON 10/100, etc)



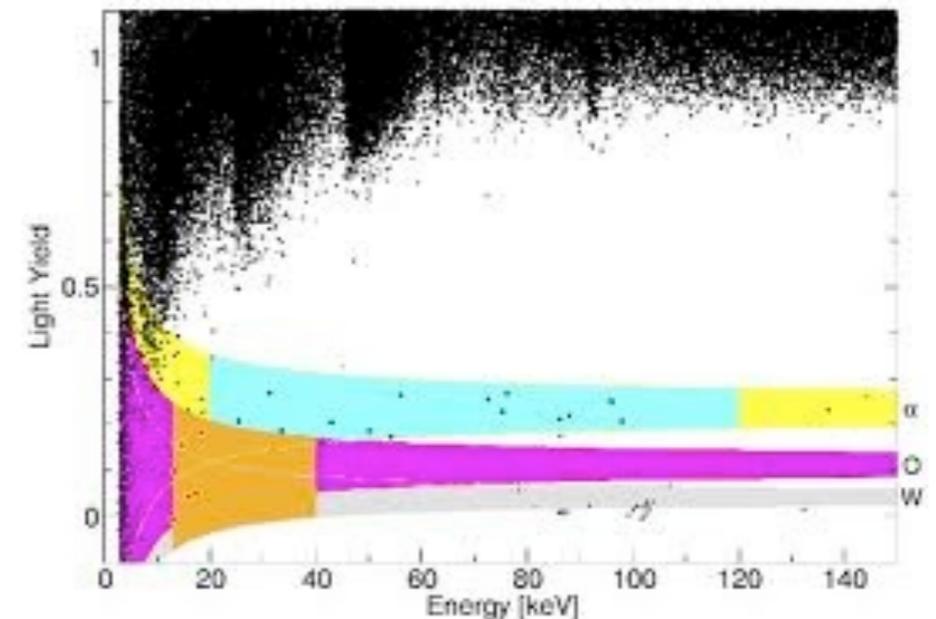
Hints for Light WIMPs

8

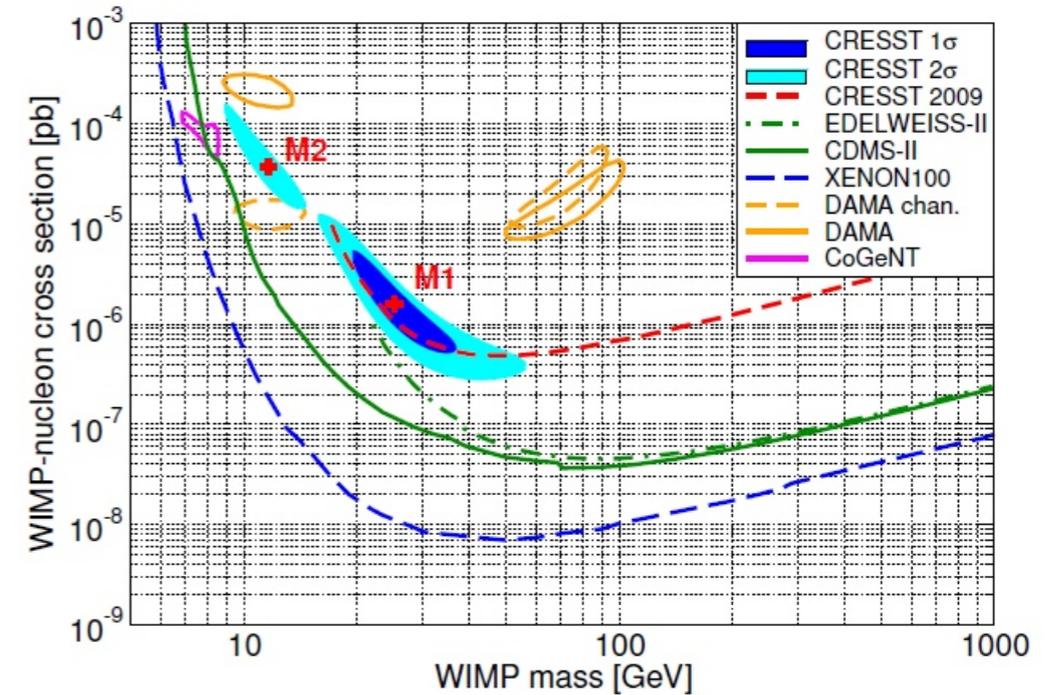
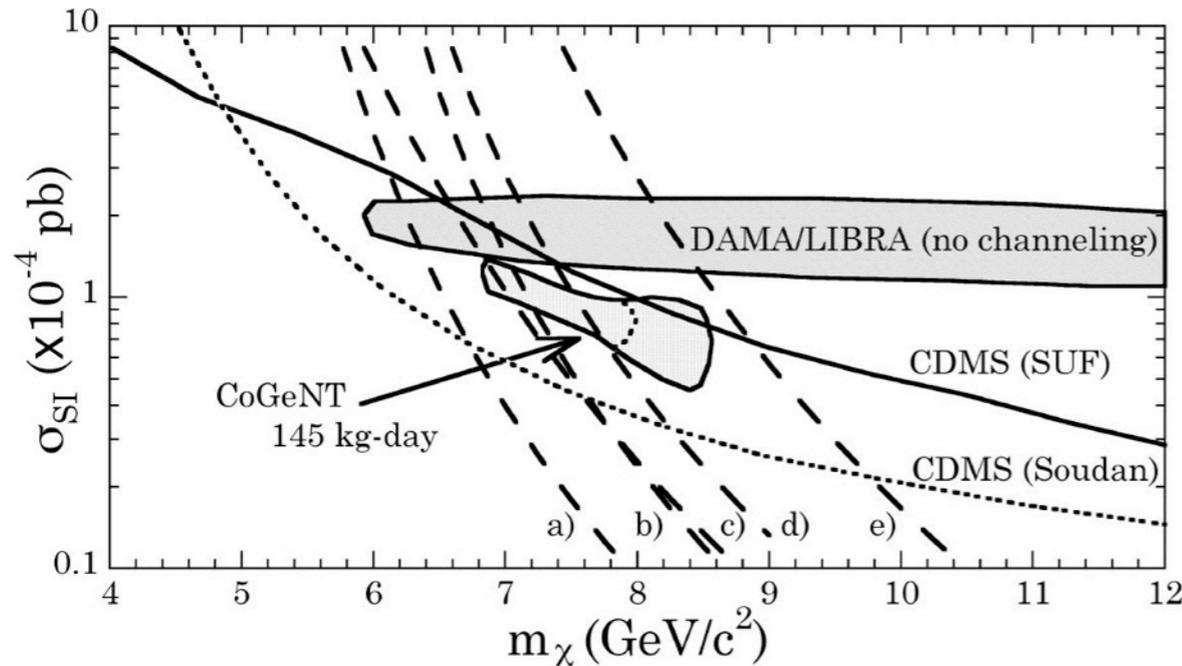
- **DAMA/LIBRA (NaI):** annual modulation effect at 8.9σ C.L.
- **CoGeNT (Ge):** excesses of events(2010), annual modulation effect at 2.8σ C.L. (2010).
- **CRESST II (CaWO₄):** excesses of events (2011).



CoGeNT



CRESST II



- Tracks of **low mass WIMPs**?
- Tension with the constraints from null expts, XENON10/100, CDMS, ...
- No simultaneous explanation of CoGeNT, DAMA and CRESST-II.

Isospin-Violating DM

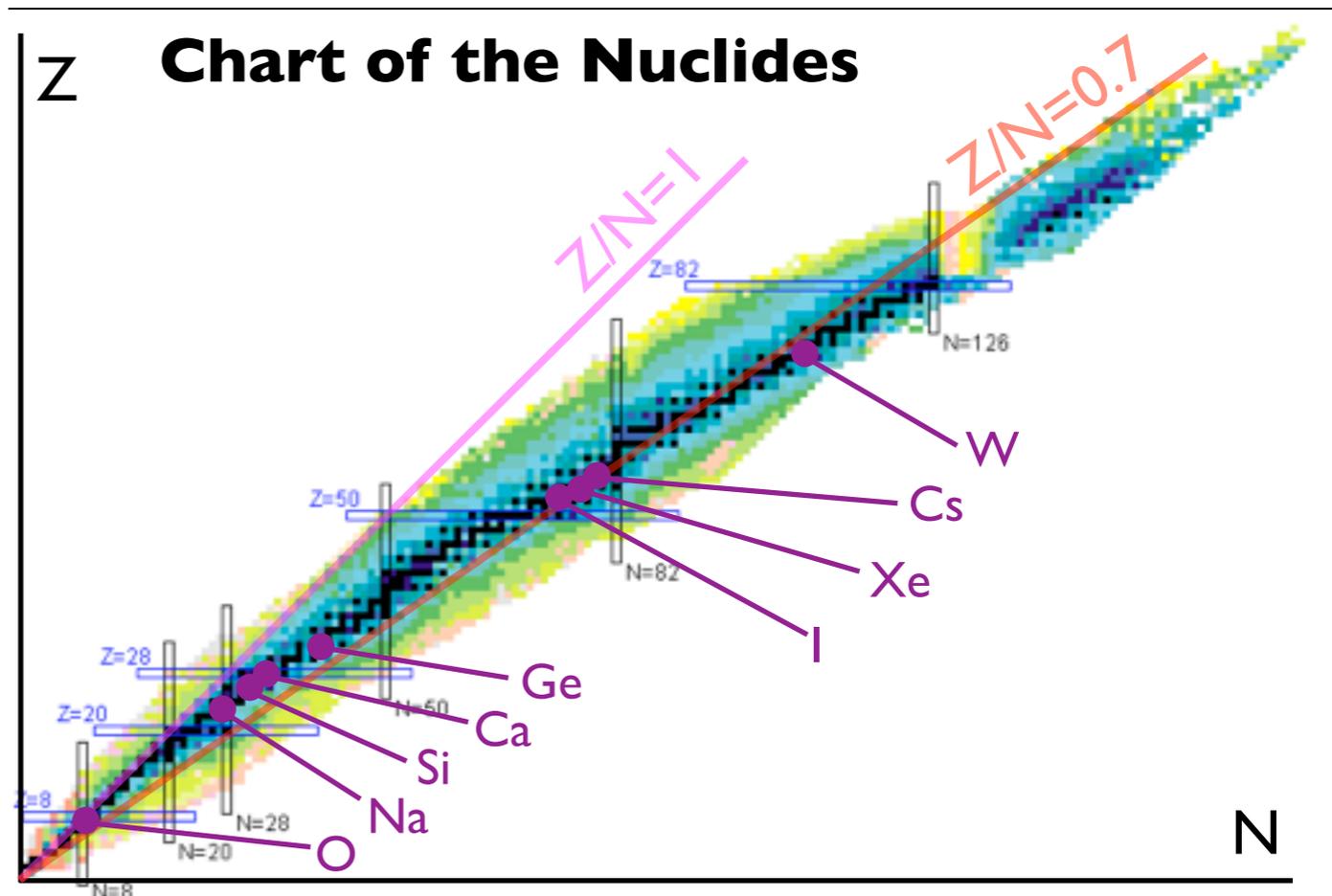
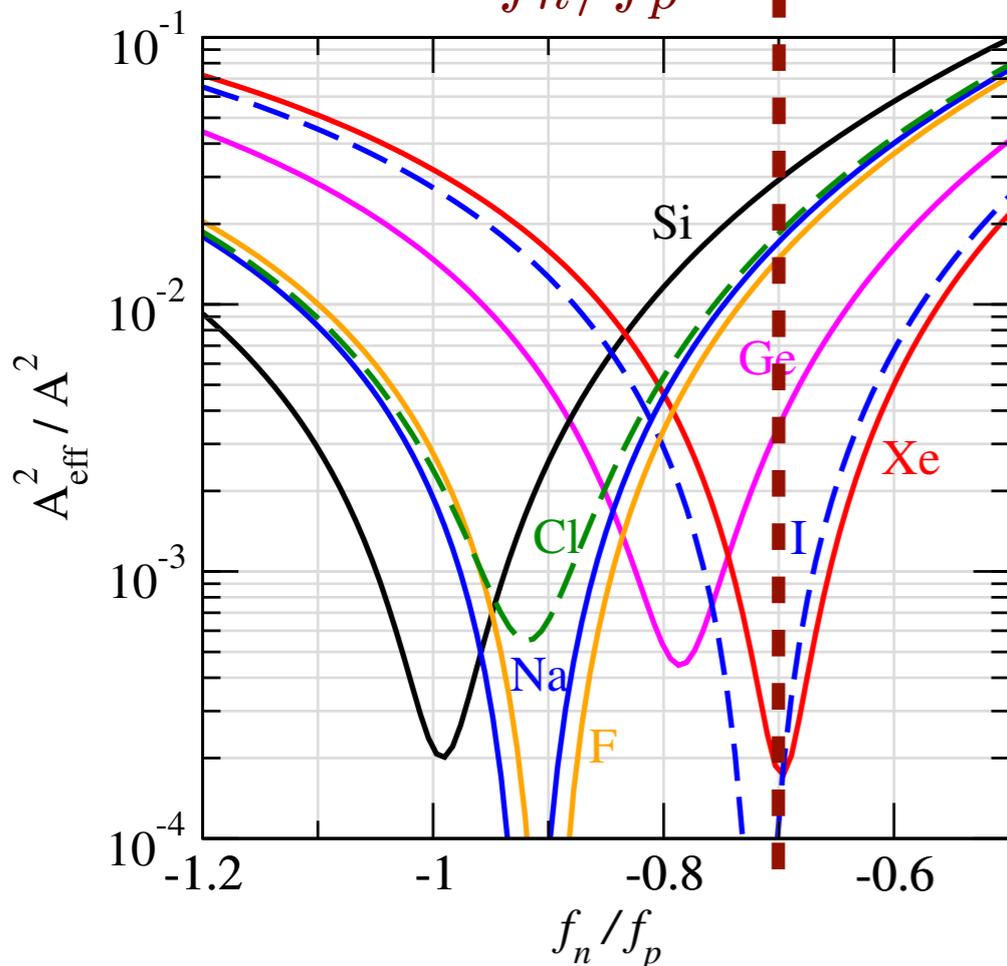
Kurylov, Kamionkowski(2004);Giuliani (2005);
 Chang, Liu, Pierce, Weiner, Yavin(2010);Feng, Kumar, Marfatia, Sanford (2011),

- DM elastic scattering off nucleus: **coherent effect!**

$$f_n = f_p \Rightarrow \sigma_{\chi N} \propto (Zf_p + (A - Z)f_n)^2 \propto A^2$$

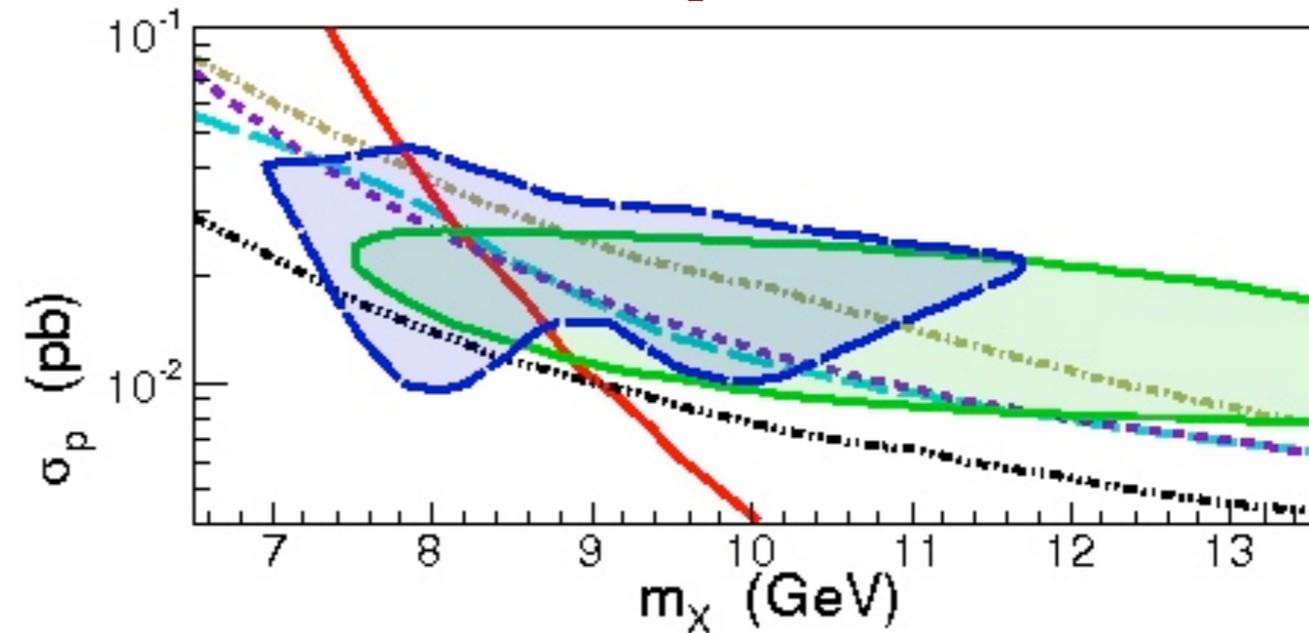
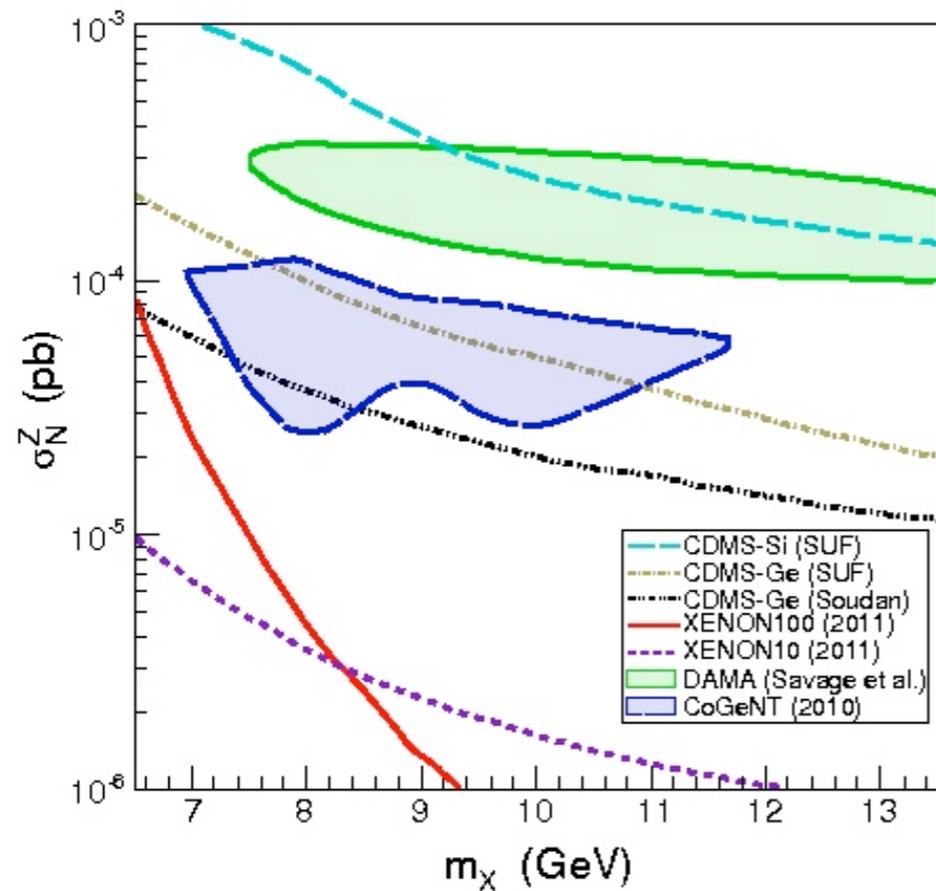
- Breaking this assumption (isospin-violation): release one additional parameter: f_n/f_p

$$f_n/f_p = -0.7$$

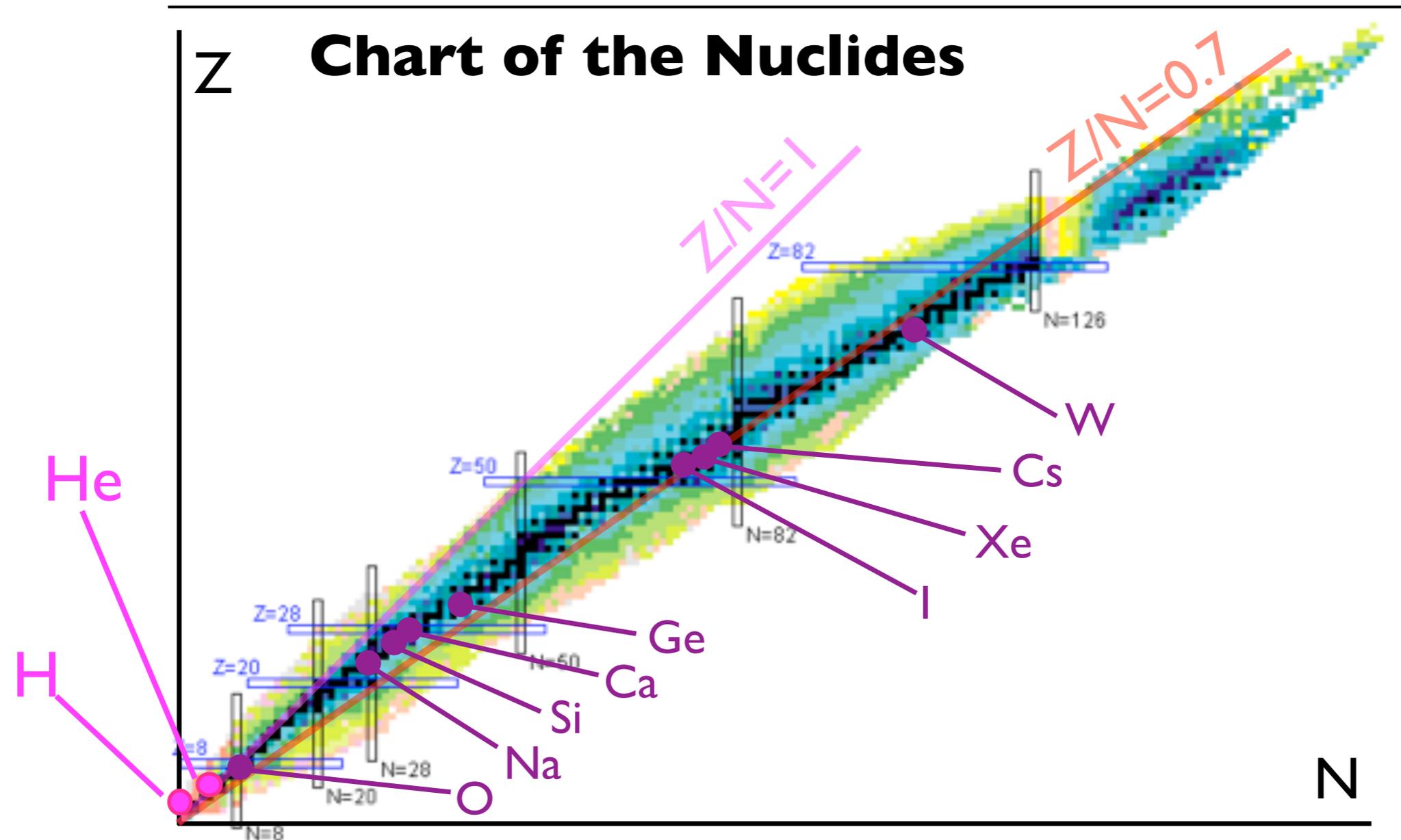


$$f_n/f_p = 1$$

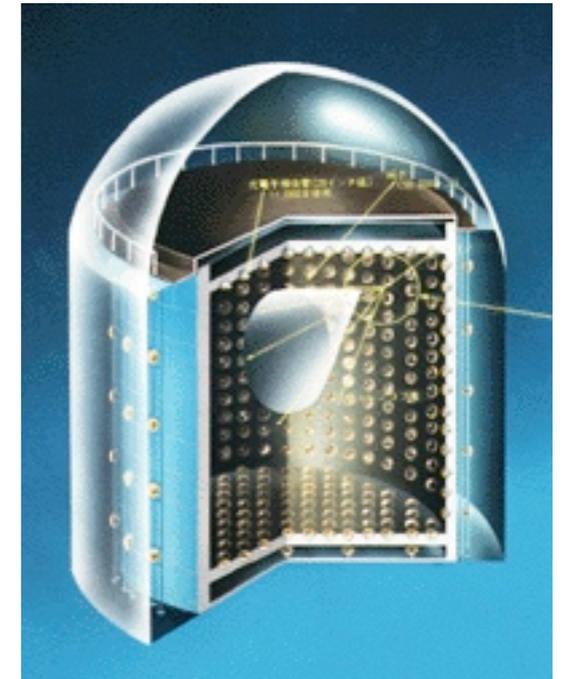
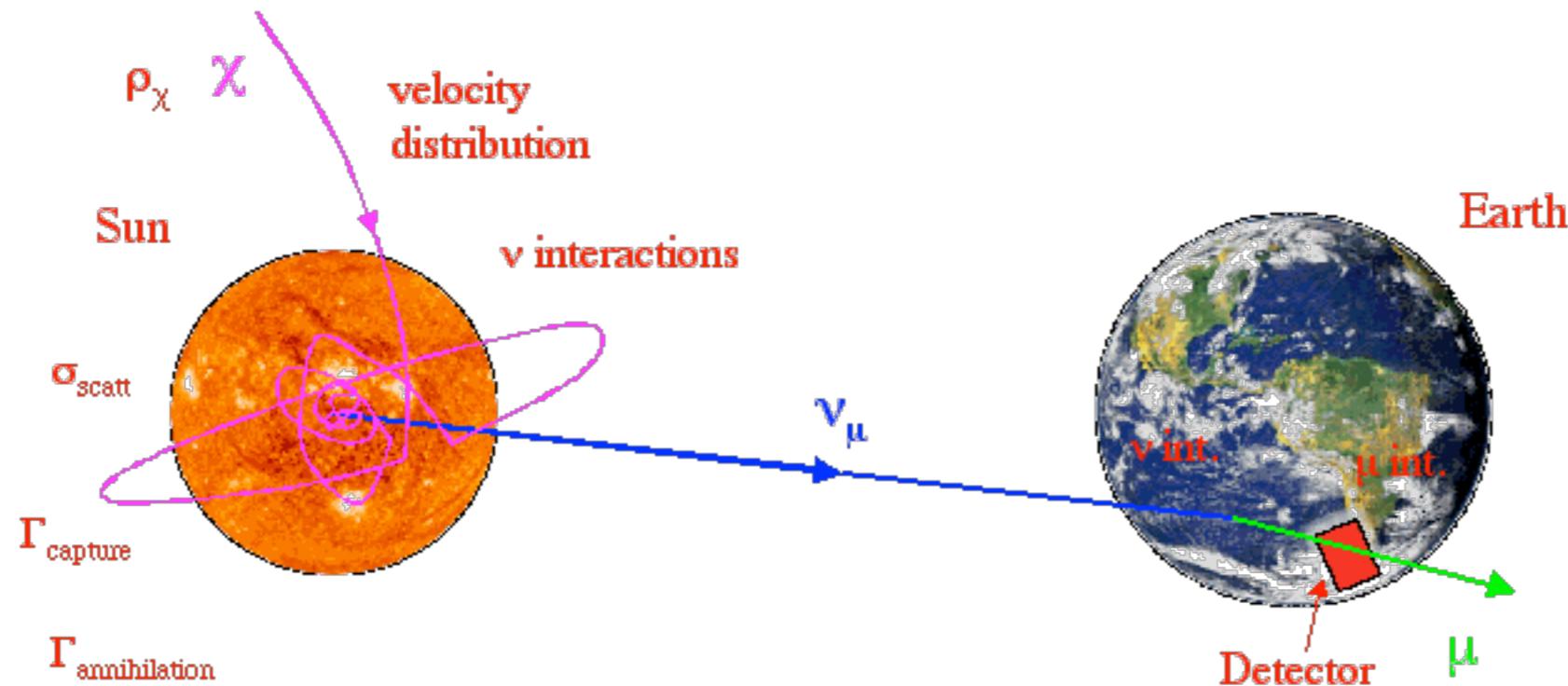
$$f_n/f_p = -0.7$$



tension alleviated



Light isotopes in the Sun: H, He, O (Z/N=1)
The Sun is a better detector for low mass IVDM!



- DM capture and self-annihilation in the Sun

$$\frac{dN}{dt} = C_\odot - C_A N^2 \longrightarrow N(t) = \sqrt{\frac{C_\odot}{C_A}} \tanh(\sqrt{C_\odot C_A} \cdot t)$$

- Capture and self-annihilation are in equilibrium

$$\Gamma_A = \frac{C_\odot}{2}$$

- The Sun is a natural detector for DM direct detection.

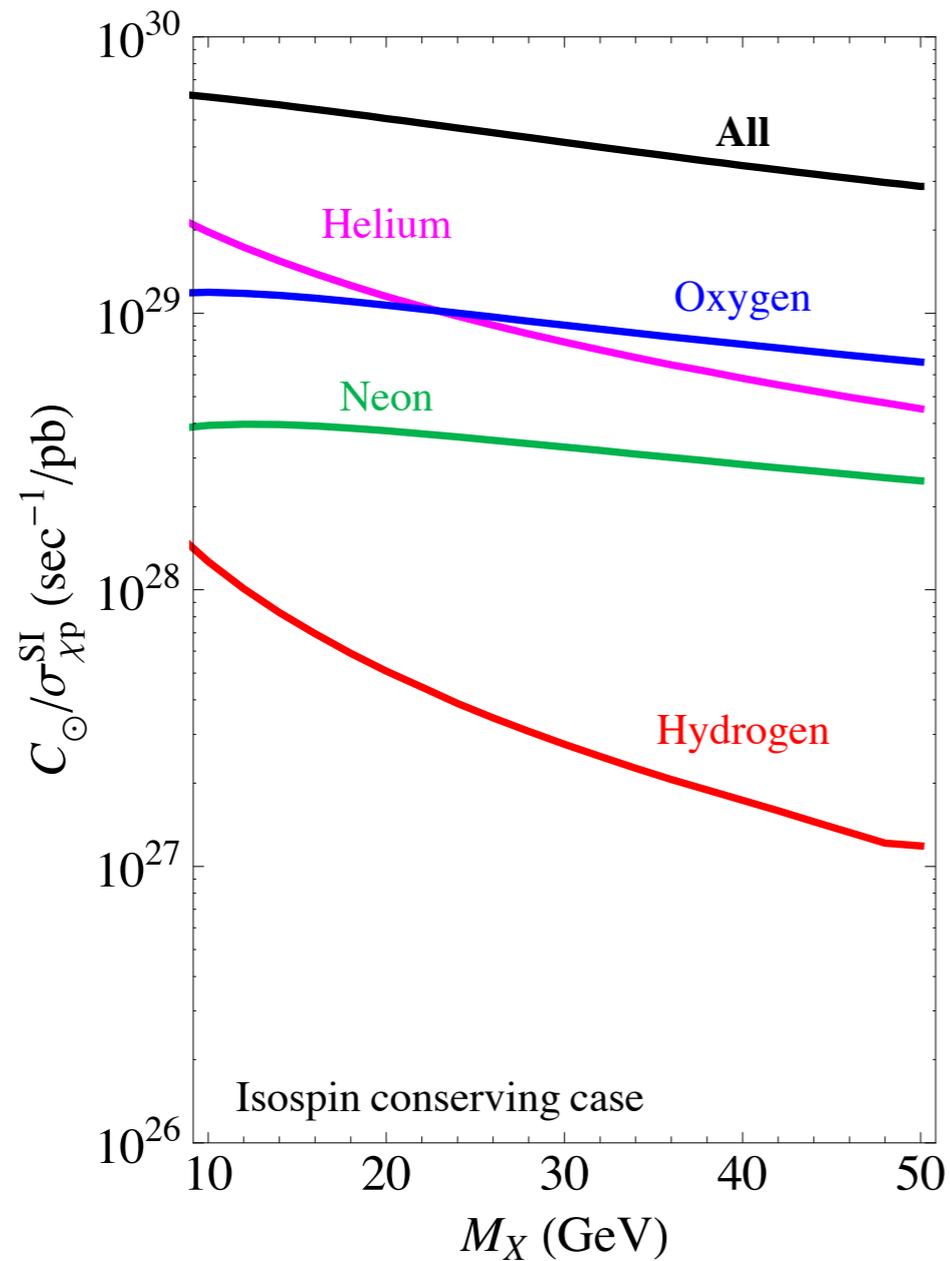
● The capture rate

$$C_{\odot,i} = \sum_i 4\pi \int_0^{R_{\odot}} r^2 dr \frac{\rho_{\chi} \rho_{\odot,i}(r)}{2m_{\chi} \mu_i^2} \sigma_i \\ \times \int_0^{\infty} du \frac{f(u)}{u} \theta(E_{R,\max} - E_{R,\text{cap}}) \int_{E_{R,\text{cap}}}^{E_{R,\max}} dE_R F^2(E_R),$$

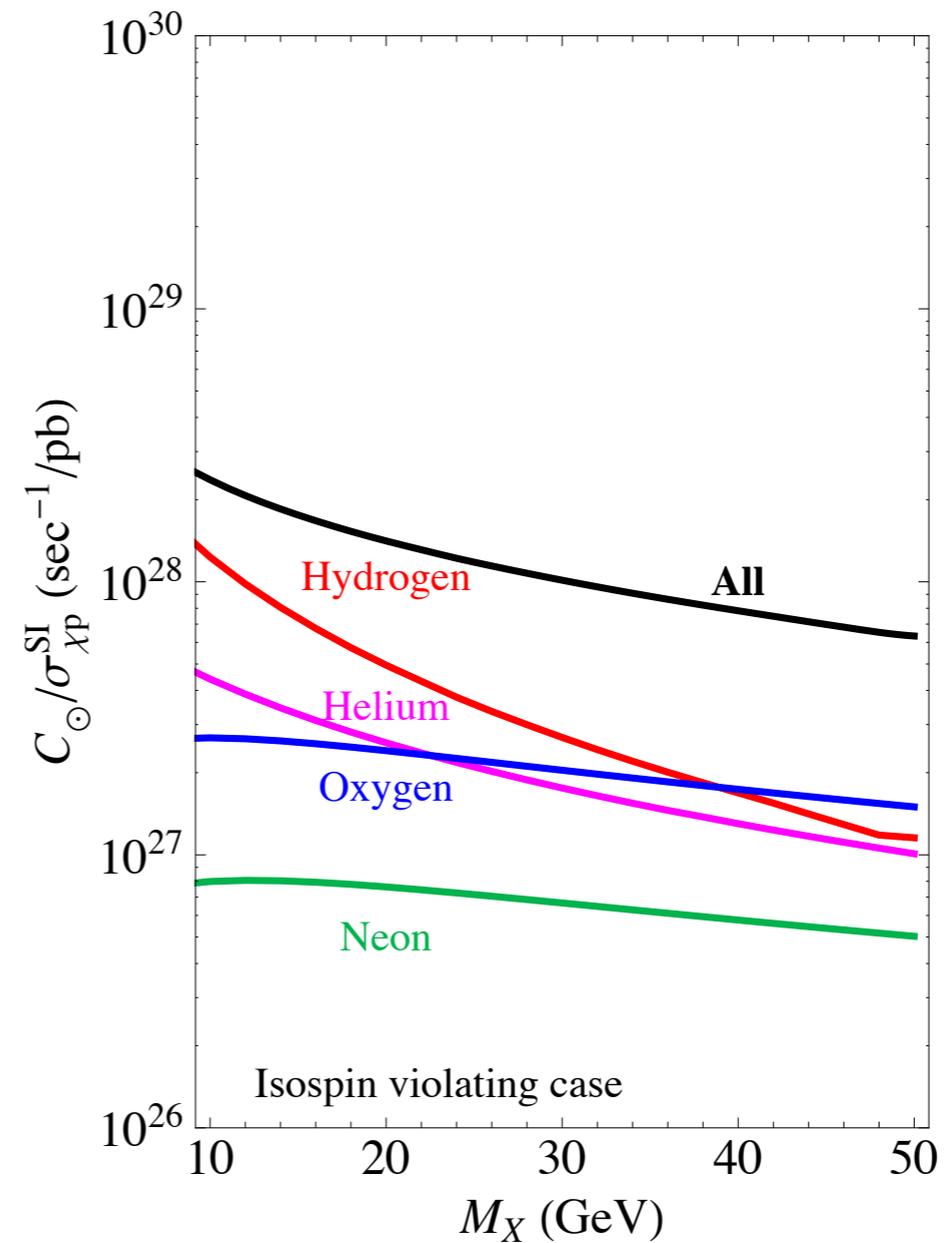
● The suppression factors

Element	Xe	Ge	Na	Solar capture
Suppression	1.3×10^{-4}	2.6×10^{-3}	1.3×10^{-2}	4.0×10^{-2}

$$f_n/f_p = -0.7 \text{ and } m_{\chi} = 10 \text{ GeV.}$$

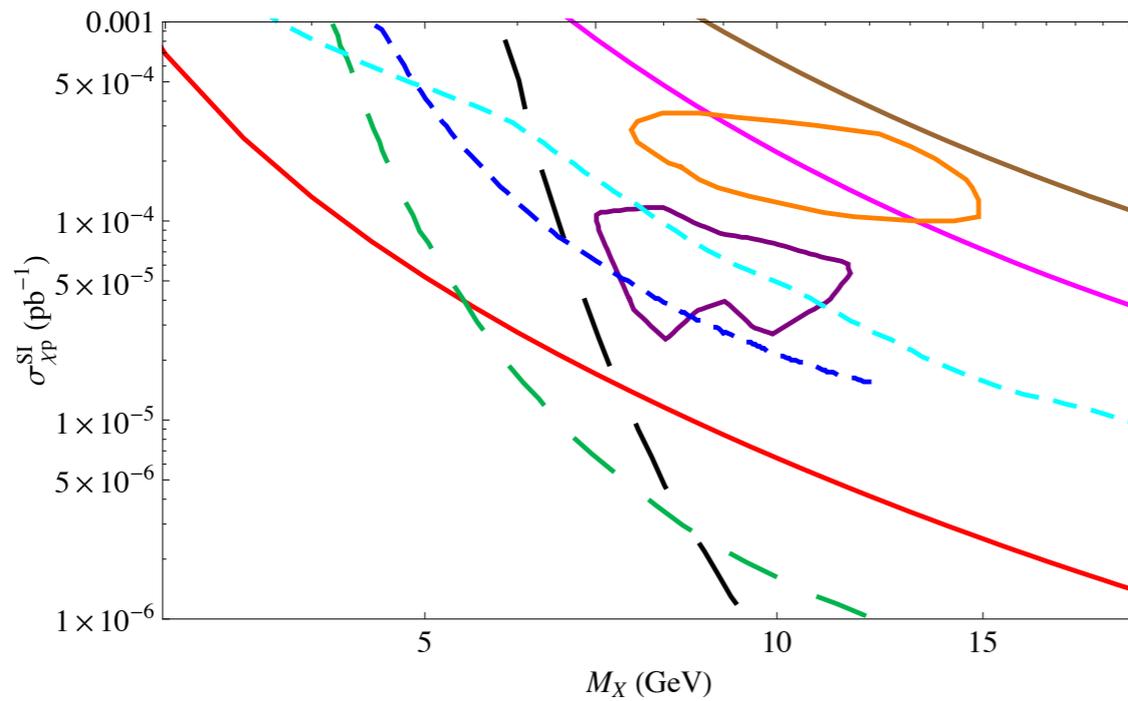


Isospin-conserving DM

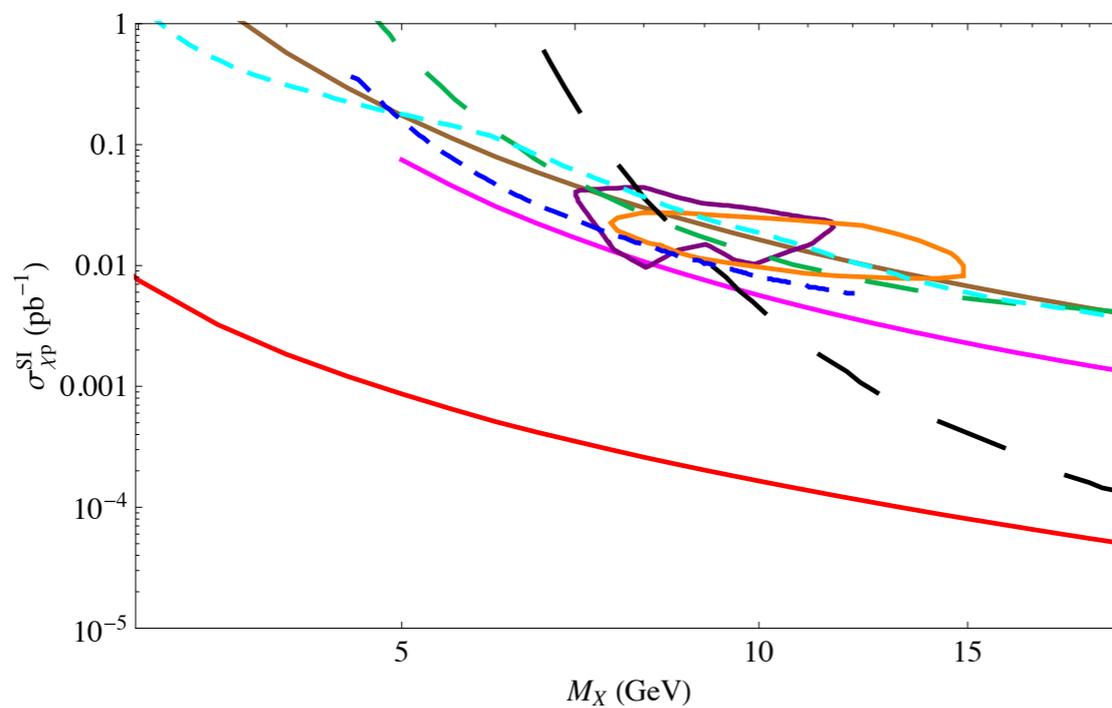


Isospin-violating DM

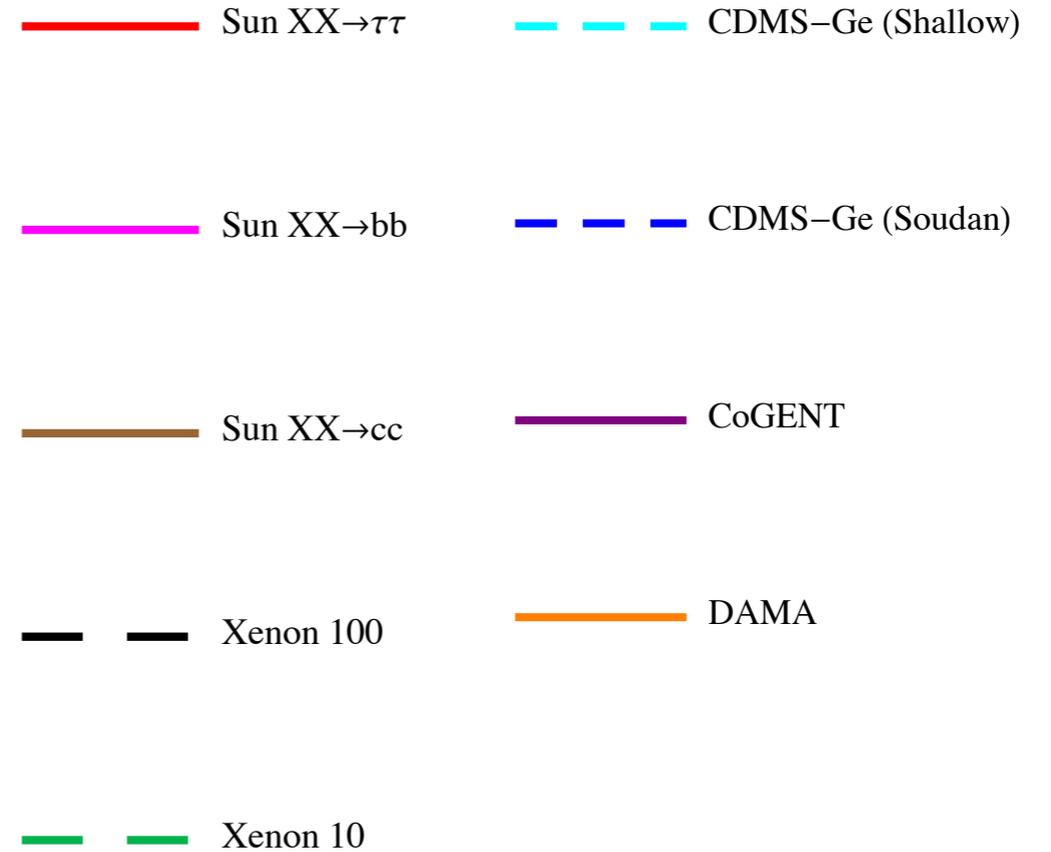
The constraints from Super-Kamiokande data 16



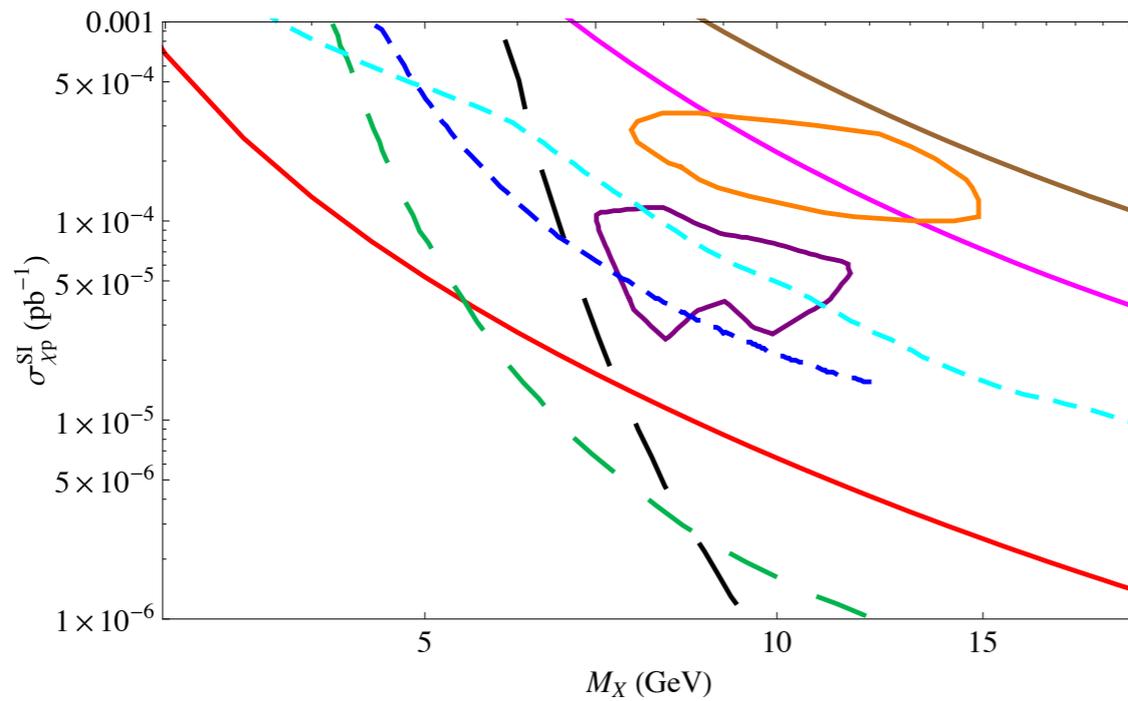
Isospin-conserving DM



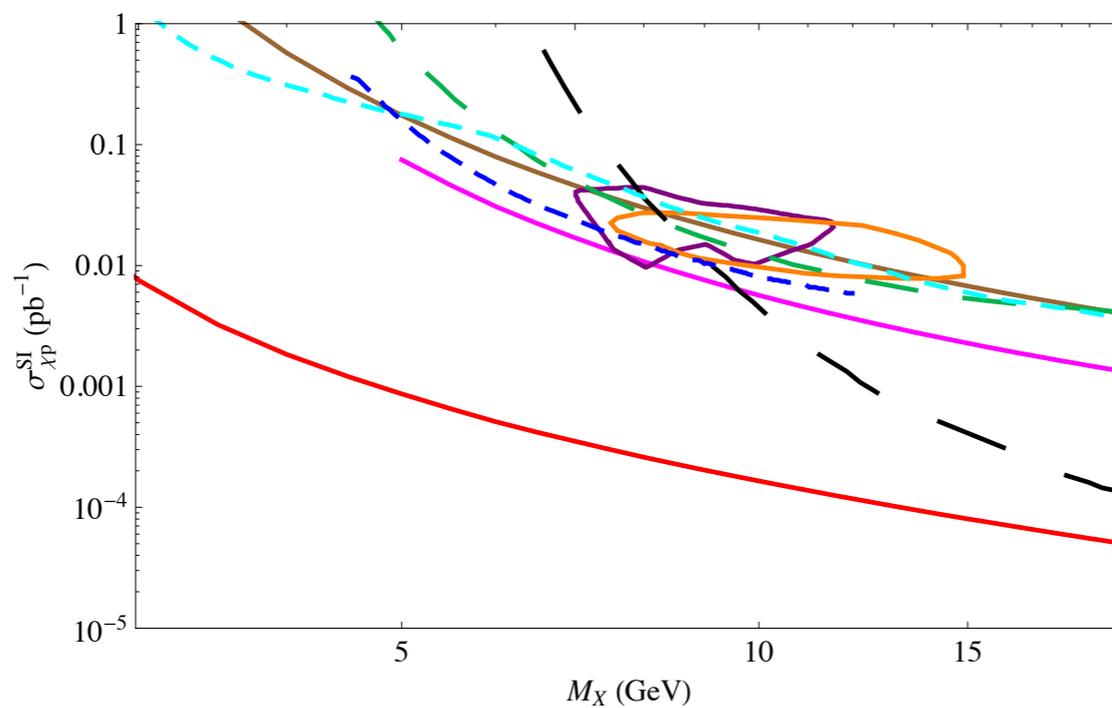
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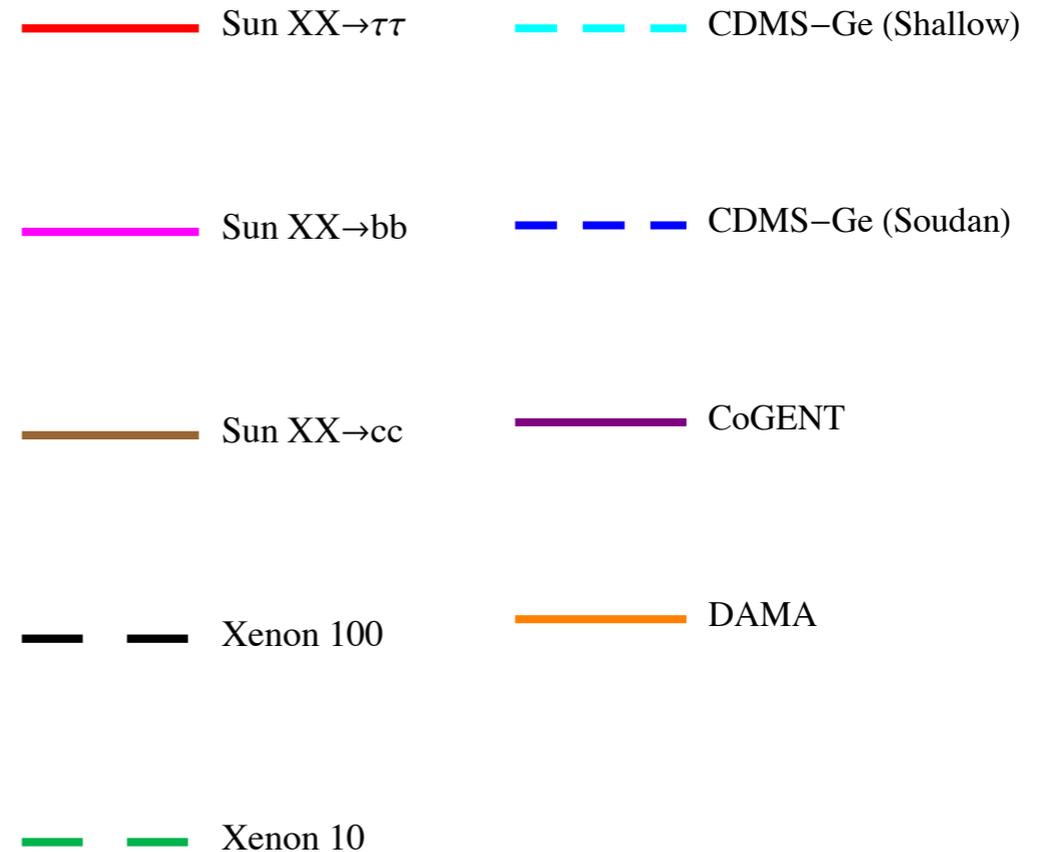
The constraints from Super-Kamiokande data 16



Isospin-conserving DM



Isospin-violating DM



Challenge to the IVDM model building

- **Low mass DM explanation for direct detection expt's: in tension with constraints**
- **IVDM: a phenomenological way proposed model is helpful to alleviate the tension.**
- **The neutrino flux from the Sun puts severe constraint to the IVDM.**