Thermal Mass Spectra of Vector and Axial-Vector Mesons in Predictive Soft-Wall AdS/QCD Model



Ling-xiao Cui 2012. 5. 10



Introduction

Gauge/Gravity Correspondence

➤ Bottom-Up Holographic QCD



The AdS/CFT Correspondence

4 dimensional N=4 Super Symmetric $SU(N_c)$ Yang-Mills theory $(N_c \to \infty)$ is dual to Type II B Super String theory in an $AdS \times S_5$ background.

J. Maldacena 1998

Gauge/Gravity Correspondence: Extend to less or non SUSY theory

- d+1 Gravity on the bulk

- d QFT on the boundary



- the classical gravity theory

- strongly coupled QFT



AdS/CFT Dictionary

Witten 98; Gubser, Klebanov, Polyakov 98

- 5D bulk field ϕ
- 5D mass m_5
- 5D gauge symmetry
- small z
- Confinement
- Kaluza-Klein states





Current (global symmetry)



(IR) cutoff z_m (in Hard Wall)

Excited, Resonant spectrum

$$\langle e^{\int d^d x \, \phi_0(\vec{x}) \mathcal{O}(\vec{x})} \rangle_{4-D \, \mathcal{N}=4 \, \text{SU(N) SYM}} = Z_{\text{Sugra}} \Big|_{\phi(0,\vec{x})=\phi_0(\vec{x})}$$

GKP-W relation



AdS/QCD

Goal: Try to understand QCD using the 5 dimensional dual gravity theory (AdS/CFT correspondence)

Approaches:

- ◆ Top-down Approach: From String Theory Find brane config. for the gravity dual D3-D7 system; D4-D8 system
- Bottom-up Approach : From phenomenological Introduce fields, etc. as needed based on the AdS/CFT
 - * Hard Wall Model Introduce IR brane for confinement
 - * Soft Wall Model dilaton running

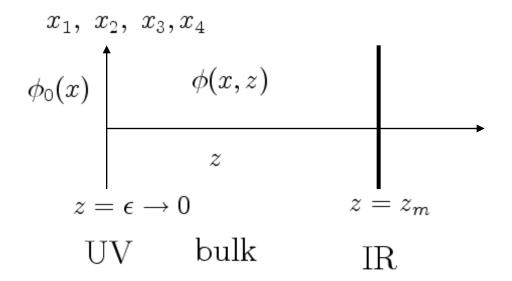


Hard wall model

J. Erlich, E. Katz, D. T. Son and M. A. Stephanov, Phys. Rev. Lett. 95 (2005)

- lacktriangle Infrared Brane at $z=z_mpprox rac{1}{\Lambda_{
 m QCD}}$
- Metric Slice of AdS metric

$$ds^{2} = \frac{1}{z^{2}}(-dz^{2} + dx^{\mu}dx_{\mu}), \qquad 0 < z \le z_{m}.$$





Hard wall model

J. Erlich, E. Katz, D. T. Son and M. A. Stephanov, Phys. Rev. Lett. 95 (2005)

Boundary

$$ar q_L \gamma^\mu t^a q_L$$

$$\bar{q}_R \gamma^\mu t^a q_R$$

$$\overline{q}_R^{lpha}q_L^{eta}$$



$$A_{L\mu}^a$$

$$A^a_{Ru}$$

$$(2/z)X^{\alpha\beta}$$

Global $SU(3)_L \times SU(3)_R$ symmetry

 $SU(3)_L \times SU(3)_R$ gauge symmetry

5D action:
$$S = \int \! d^5 x \, \sqrt{g} \, \operatorname{Tr} \Bigl\{ |DX|^2 + 3 |X|^2 - \frac{1}{4g_5^2} (F_L^2 + F_R^2) \Bigr\}$$

Mass term is determined by the scaling dimension

$$(\Delta-p)(\Delta+p-4)=m_5^2$$

Xij has dimension $\triangle = 3$ and form p=0, AL & AR have dimension $\triangle = 3$ and form p=1



Gauge coupling

In AdS, correlation function of source fields on UV brane

$$\langle J_{\mu}J_{\nu}\rangle = \frac{\delta}{\delta V_{\mu}} \frac{\delta}{\delta V_{\nu}} e^{iS} = \frac{1}{g_5^2} V_{\mu} \frac{1}{z} \partial_z V(z) \Big|_{z=\epsilon}$$

Bulk to boundary propagator solution to the equation of motion with V(0)=1

$$V(Q,z) = 1 + \frac{Q^2 z^2}{4} \ln(Q^2 z^2) + \dots$$
 $\Pi_V(Q^2) = -\frac{1}{2a_z^2} \ln Q^2$



$$\Pi_{
m V}(Q^2) = -rac{1}{2g_5^2} \ln Q^2$$

In QCD, correlation function of vector current is

$$\int d^4x \langle J^a_{\mu}(x) J^b_{\nu}(0) \rangle = \delta^{ab} \left(q_{\mu} q_{\nu} - q^2 g_{\mu\nu} \right) \Pi_V(Q^2) \qquad \Pi_V(Q^2) = -\frac{N_c}{24\pi^2} \ln(Q^2)$$

$$\Pi_V(Q^2) = -\frac{N_c}{24\pi^2} \ln(Q^2)$$

$$\left(g_5^2=rac{12\pi^2}{N_c}
ight)$$



small coupling



Chiral symmetry breaking by X

One can find the expectation value of *X*

$$X_0(z) = \frac{1}{2}Mz + \frac{1}{2}\Sigma z^3$$

defined as the classical solutions of field equations, using UV boundary conditions

$$(2/\epsilon)X(\epsilon) = M$$

M is quark mass matrix (Explicit chiral breaking)

Σ is chiral condensate (Spontaneous chiral breaking)

the model has three free parameters: M, Σ , Zm, which are fixed from the experiment data. (ρ meson, pion mass and pion decay constant)



Soft wall model

Introducting a background dilaton $\Phi(z \to \infty) = \mu_d^2 z^2$ can lead to a linear

trajectory for resonance vector meson mass

$$I = \int d^5x \, e^{-\Phi(z)} \sqrt{g} \left\{ -|DX|^2 + 3|X|^2 - \frac{1}{4g_5^2} (F_L^2 + F_R^2) \right\}$$

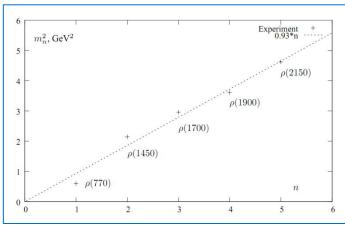
Solving equations of motion for vector field

$$-\psi_n'' + V(z)\psi_n = m_n^2 \psi_n \qquad V(z) = \frac{1}{4} (B')^2 - \frac{1}{2} B''$$

$$B = \Phi - A = z^{2} + \log z \qquad \psi_{n}(z) = e^{-\frac{z^{2}}{2}} z^{m+1/2} \sqrt{\frac{2n!}{(m+n)!}} L_{n}^{m}(z^{2})$$

Linear trajectory for mass spectra of vector mesons

$$m_n^2 = 4(n+1)$$





Modified Soft-Wall AdS/QCD

Y.Q.Sui, YLW, Z.F.Xie, Y.B.Yang PRD arXiv:0909.3887

- Hard-wall AdS/QCD models contain the chiral symmetry breaking, the resulting mass spectra for the excited mesons are contrary to the experimental data
- Soft-wall AdS/QCD models describe the linear confinement and desired mass spectra for the excited vector mesons, while the chiral symmetry breaking can't consistently be realized.
- A quartic interaction in the bulk scalar potential was introduced to incorporate linear confinement and chiral symmetry breaking. While it causes an instability of the scalar potential and a negative mass for the lowest lying scalar meson state.
- How to naturally incorporate two important features into a single AdS/QCD model and obtain the consistent mass spectra.



Modified Soft-Wall AdS/QCD

The 5D action with the background field of dilaton $\Phi(z)$ and a quartic term in the bulk scalar potential

$$S_5 = \int d^5x \sqrt{g} e^{-\Phi(z)} \operatorname{Tr} \left[|DX|^2 - m_X^2 |X|^2 - \lambda |X|^4 - \frac{1}{4g_5^2} \left(F_L^2 + F_R^2 \right) \right]$$

Deformed 5D Metric in IR Region

$$ds^{2} = a^{2}(z) \left(\eta_{\mu\nu} dx^{\mu} dx^{\nu} - dz^{2} \right)$$
$$a^{2}(z) = (1 + \mu_{g}^{2} z^{2})/z^{2}$$

- IR region is modified
- UV region is not changed

Equation for the bulk vacuum:

$$\partial_z \left(a^3(z) e^{-\Phi} \partial_z v(z) \right) - a^5(z) e^{-\Phi} m_X^2 v(z) = 0$$

IR&UV behaviour of dilaton and vacuum:

$$v(z \to \infty) = \gamma \, (\mu_d z)^{\alpha}$$

$$v(z \to 0) = m_q \zeta z + \frac{\sigma z^3}{\zeta}$$

Solutions for the dilaton field at the UV & IR boundary

$$\Phi(z \to 0) = 3\mu_g^2 z^2 + O(z^4)$$

$$\Phi(z\to\infty) = \mu_d^2 \, z^2$$



Three type of models for v(z)

Models	v(z)	参数	
Ia	$z(A+Bz^2)(1+Cz^2)^{-1}$	$B = \frac{\sigma}{\zeta} + m_q \zeta C,$	$C = B/\mu_d \gamma$
Ib	$z(A+Bz^2)(1+Cz^2)^{-5/4}$	$B = \frac{\sigma}{\zeta} + \frac{5}{4} m_q \zeta C,$	$C = (B^2/\mu_d \gamma^2)^{2/5}$
IIa	$z(A+Bz^2)(1+Cz^4)^{-1/2}$	$B = \frac{\sigma}{\zeta},$	$C = (B/\mu_d \gamma)^2$
IIb	$z(A+Bz^2)(1+Cz^4)^{-5/8}$	$B = \frac{\sigma}{\zeta},$	$C = (B^2/\mu_d \gamma^2)^{4/5}$
IIIa	$z[A+B\tanh(Cz^2)]$	$B = \mu_d \gamma - m_q \zeta,$	$C = \frac{\sigma}{\zeta B}$
IIIb	$z[A + B \tanh(Cz^2)](1 + Gz^4)^{-1/8}$	$B = \mu_d^{1/2} \gamma G^{1/8} - m_d$	$_{q}\zeta, C = \frac{\sigma}{\zeta B}$

$$A = m_q \zeta$$

Two IR boundary conditions of the bulk VEV:

Case a:
$$v(z \to \infty) = \gamma(\mu_d z)$$

Case b:
$$v(z \to \infty) = \gamma(\sqrt{\mu_d z})$$



参数输入

 $igoplus \mu_g$ scales the mass spectra of meson resonances, it can be found out from a global fitting.

$$\mu_{\rm g} = 363 \, {\rm MeV}$$
 (Case a) $\mu_{\rm g} = 257 \, {\rm MeV}$ (Case b)

lackloais The three parameters m_q , σ and γ are fixed by the known experimental values:

$$m_{\pi} = 139.6 \text{ MeV}$$
 $f_{\pi} = 92.4 \text{ MeV}$ $f_{\pi}^2 m_{\pi}^2 = 2m_q \sigma$

Parameter	Ia	Ib	IIa	IIb	IIIa	IIIb
$m_q \; ({\rm MeV})$						
$\sigma^{\frac{1}{3}}$ (MeV)	275	265	265	272	255	268
γ	0.178	0.136	0.153	0.112	0.164	0.112

for $\lambda = 0$



Solving Equations of Motion

Pseudoscalar Sector

$$\partial_z (a(z)e^{-\Phi}\partial_z \phi^a) + g_5^2 a^3(z) v^2(z)e^{-\Phi} (\pi^a - \phi^a) = 0$$
$$q^2 \partial_z \phi^a - g_5^2 a^2(z) v^2(z) \partial_z \pi^a = 0$$

Scalar Sector

$$\partial_z \left(a^3(z) e^{-\Phi} \partial_z S_n(z) \right) - a^5(z) e^{-\Phi} m_X^2 S_n(z) = -a^3(z) e^{-\Phi} m_{S_n}^2 S_n(z)$$

Vector Sector

$$-\partial_z^2 V_n + \omega' \partial_z V_n = m_{V_n}^2 V_n$$

Axial-vector Sector

$$e^{\Phi}\partial_z(a(z)e^{-\Phi}\partial_z A_n) + a(z)q^2 A_n - a^3(z)g_5^2 v^2(z)A_n = 0$$

IR & UV Boundary Condition:

$$s_n(z \to 0) = 0$$
, $\partial_z s_n(z \to \infty) = 0$,



Mass Spectra of Vector Mesons

Without Quartic Interaction of bulk scalar

$$\lambda = 0$$

n	ho实验值(MeV)	Ia	Ib	IIa	IIb	IIIa	IIIb
0	775.5 ± 1	739	603	777	727	775	748
1	1465 ± 25	1223	1175	1292	1468	1303	1501
2	1720 ± 20	1534	1509	1596	1744	1610	1773
3	1909 ± 30	1784	1769	1842	1971	1856	1999
4	2149 ± 17	2000	1990	2054	2170	2068	2196
5	2265 ± 40	2193	2187	2249	2351	2255	2373
6		2370	2367	2417	2516	2426	2535
7		2534	2532	2578	2671	2584	2685

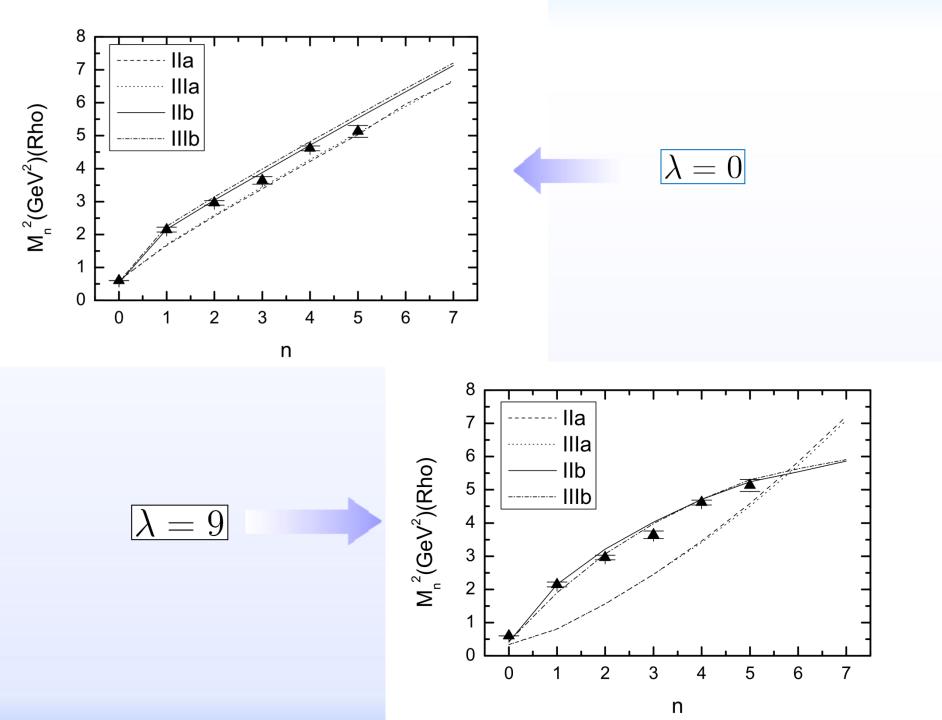


Mass Spectra of Vector Mesons

With Quartic Interaction of bulk scalar

$$\lambda = 9$$

n	ho实验值(MeV)	IIa	IIb	IIIa	IIIb
0	775.5 ± 1	583	646	584	661
1	1465 ± 25	900	1468	893	1378
2	1720 ± 20	1248	1793	1252	1753
3	1909 ± 30	1564	2008	1565	1994
4	2149 ± 17	1860	2170	1847	2174
5	2265 ± 40	2143	2289	2122	2332
6		2417	2353	2395	2372
7		2685	2420	2661	2432





For finite temperature



deconfinement phase transition

gravity theory

Hawking-Page transition

[Herzog , Phys.Rev.Lett.98:091601,2007]

• AdS BH (Deconfinement)
$$ds^2 = \frac{L^2}{z^2} \left(f(z) dt^2 + d\vec{x}^2 + \frac{dz^2}{f(z)} \right)$$

$$f(z) = 1 - (z/z_h)^4 \qquad z_h \text{:Horizon}$$

dual

(HP) Hawking temperature $T = 1/(\pi z_h)$ Temperature of the medium

• Thermal AdS (Confinement) $ds^2 = L^2 \left(\frac{dt^2 + d\vec{x}^2 + dz^2}{z^2} \right)$



Phase Transition

C.P.Herzog, PRL 98.091601 (2007)

On-shell action density



Free energy

$$I=-\frac{1}{2\kappa^2}\int d^5x\sqrt{g}\left(R+\frac{12}{L^2}\right) \qquad \text{On-shell} \qquad I=\frac{4}{L^2\kappa^2}\int d^5x\sqrt{g}$$

On-shell
$$I = \frac{4}{L^2 \kappa^2} \int d^5 x \sqrt{g}$$

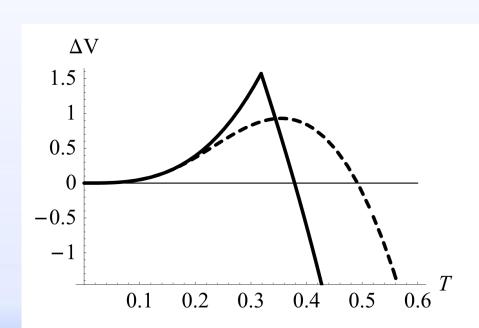
Difference between V1 and V2

$$\Delta V = \lim_{\epsilon \to 0} \left(V_2(\epsilon) - V_1(\epsilon) \right)$$

When $\triangle V$ is positive (negative), tAdS (AdSBH) is stable.

Critical Temperature (1st order)

Tc,Hard = 122 and Tc,soft = 191 MeV.





Extend to finite temperature

L.X.Cui, Shingo Takeuchi, Y.L.Wu arXiv:1112.5923

5D action:

$$S_5 = \int d^5x \sqrt{g} e^{-\Phi(z)} \operatorname{Tr} \left[|DX|^2 - m_X^2 |X|^2 - \lambda |X|^4 - \frac{1}{4g_5^2} \left(F_L^2 + F_R^2 \right) \right]$$

Deformed AdS Black Hole:
$$ds^2 = a^2(z) \left(f(z) dt^2 - \sum_{i=1}^3 dx_i^2 - \frac{dz^2}{f(z)} \right)$$

with
$$a^2(z)=1/z^2+\underline{\mu_g^2}$$
 • IR region is improved • If $\mu_{\rm g}=0$, it become an AdS BH

$$f(z) = 1 - (z/z_0)^4$$

 z_0 : Horizon

Hawking temperature is not changed:

$$T = 1/(z_0 \pi)$$

The EOM for Vector and Axial-Vector

Action at the quadratic order:

$$S = \int dx^{5} e^{-\Phi(z)} \sqrt{g} \operatorname{Tr} \left[-\frac{1}{2g_{5}^{2}} \left(F^{V} F^{V} + F^{A} F^{A} \right) - (2X_{0} \partial \pi + [V, X_{0}] + \{A, X_{0}\}) \left(2X_{0} \partial \pi - [V, X_{0}] + \{A, X_{0}\} \right) \right]$$

$$\text{Vector: } 0 = \widetilde{V}_x^{a\prime\prime}(p,u) + \left(\frac{a'(u)}{a(u)} + \frac{f'(u)}{f(u)} - \Phi'(u)\right) \widetilde{V}_x^{a\prime}(p,u) + \frac{1}{(\pi T)^2} \left(\frac{\omega^2}{f^2(u)} - \frac{q^2}{f(u)}\right) \widetilde{V}_x^{a}(p,u)$$

$$\begin{aligned} \textbf{Axial-Vector:} 0 &= \widetilde{W}_x^{a\prime\prime}(p,u) + \left(\frac{a'(u)}{a(u)} + \frac{f'(u)}{f(u)} - \Phi'(u)\right) \widetilde{W}_x^{a\prime}(p,u) + \frac{1}{(\pi T)^2} \left(\frac{\omega^2}{f^2(u)} - \frac{q^2}{f(u)}\right) \widetilde{W}_x^a(p,u) \\ &+ g_5^2 \frac{v^2(u)}{u^2 f(u)} \widetilde{W}_x^a(p,u). \end{aligned}$$



Taking in-falling boundary condition

From equation of V(z):

$$\Phi'(z) = \frac{3a'(z)}{a(z)} + \frac{(f(z)v'(z))'}{f(z)v'(z)} - \frac{a^2(z)}{f(z)v'(z)}m_X^2v(z)$$

We assume:

$$v(z) = v_1(z) \ln f(z) + v_0(z)$$

divergent term can be read:

$$\Phi'(z) = \frac{\ln f(z)}{v_1(z)f'(z)} \left\{ -a^2(z)v_1(z)m_X^2 + f'(z)v_1'(z) \right\} + \cdots$$

finite temperature part of vacuum:

$$v_1(z) = c_v \exp\left[\frac{m_X^2}{8(\pi T)^4 z^4} \left(\frac{1}{2} + \mu_g^2 z^2\right)\right]$$



Retarded Green's function

◆ 2 independent solutions near the boundary

$$V(u o 0)$$

$$\Phi_0^a(\omega, q, u = \epsilon) = \epsilon \ Y_1(\frac{\sqrt{\omega^2 - q^2}}{\pi T} \epsilon)$$

$$\Phi_1^a(\omega, q, u = \epsilon) = \epsilon \ J_1(\frac{\sqrt{\omega^2 - q^2}}{\pi T} \epsilon)$$

◆ 2 independent solutions near the horizon

$$V(u \to 1) \qquad V_{\pm} = (1 - u)^{\pm i \frac{\omega}{4\pi T}}$$

 V_+ : Out-coming from Black Hole \longrightarrow Advanced Green's function

V: In-falling into Black Hole ——— Retarded Green's function

[D.T.Son and A.O.Starinets (2002)]



Spectral Function

Linear combination of the 2 solution ——— In-falling solution

$$V(u) = A(\omega, q) \Phi_1(u) + B(\omega, q) \Phi_2(u) \xrightarrow{u \longrightarrow 1} V_{-}(u)$$

boundary condition: $\Phi_1 \approx 1$

$$\Phi_2 \approx \epsilon^2$$

Retarded Green's function

$$D^{R}(\omega,q) = \lim_{\epsilon \to 0} \frac{\delta^{2}}{\delta V_{0}(p) \delta V_{0}(p)} \exp(-S[V(p,\epsilon)]) \bigg|_{V_{0}=0}$$
 :GKP-W relation
$$\propto \lim(\frac{1}{\epsilon}V(\epsilon)^{*}\partial_{\epsilon}V(\epsilon))$$

$$\propto -2\left[\frac{B^{a}(\omega,q)}{A^{a}(\omega,q)} + \frac{q^{2}-\omega^{2}}{2}\left\{\gamma_{E} + \log\left(\frac{\epsilon}{2}\sqrt{\omega^{2}-q^{2}}\right)\right\}\right]$$

Spectral Function

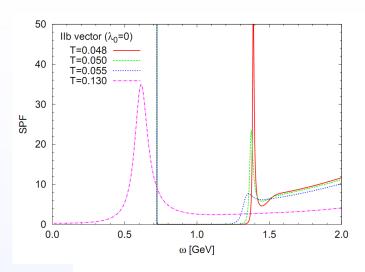
$$ho(\omega,q) = -rac{1}{\pi} {
m Im}\, G^R(\omega,q)\, heta(\omega^2-q^2) \!\propto\! rac{B^a(\omega,q)}{A^a(\omega,q)}$$

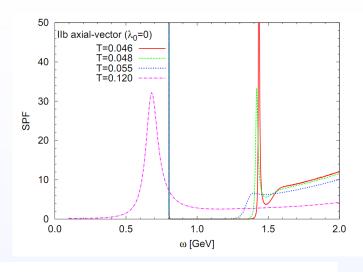


Numerical Results

◆Low temperature results







n	$\rho_{\rm V}~({\rm MeV})$	Ia	Ib	IIa	IIb	$\rho_{\rm AV}~({ m MeV})$	Ia	Ib	IIa	IIb
0	775.5 ± 1	739	603	777	727	1230 ± 40	934	714	940	807
1	1465 ± 25	1223	1175	1292	1468	1647 ± 22	1468	1247	1496	1507
2	1720 ± 20	1534	1509	1596	1744	1930^{+30}_{-70}	1822	1573	1831	1778
3	1909 ± 30	1784	1769	1842	1971	2096 ± 122	2109	1829	2102	2003
4	2149 ± 17	2000	1990	2054	2170	2270^{+55}_{-40}	2358	2049	2338	2202
5	-	2193	2187	2249	2351	-	2582	2049	2338	2202

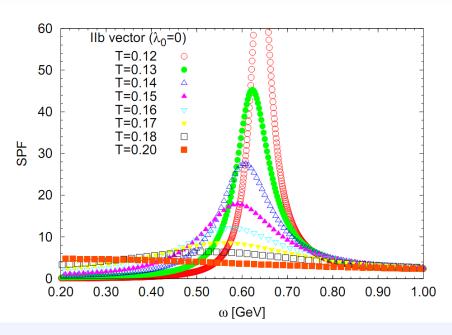
At low temperature, sharp peaks stand in accord with t=0 spectrum.

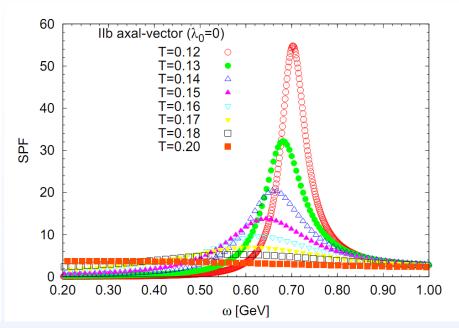


Numerical Results

Higher temperature results







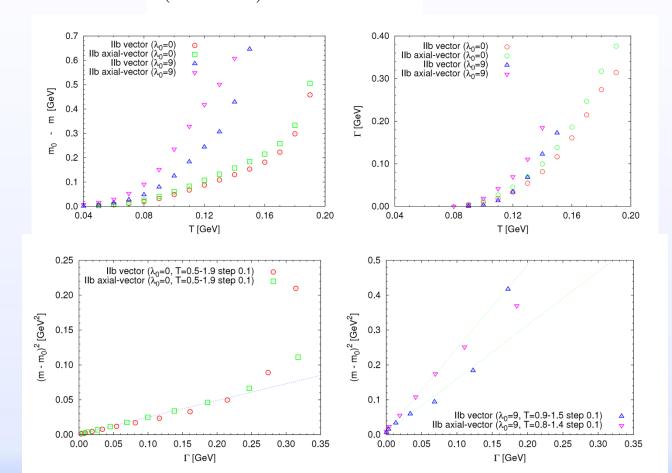
Properties

- 1. The lowest-lying state melts gradually as T increases.
- 2. The peak moves to a smaller mass with increasing T.
- 3. The excited states melt much earlier and shift more.
- 4. The lowest-lying state survies until T~200MeV

Mass shift and width of peak from spectral function

• fitting the spectral function by following Breit-Wigner form:

$$rac{a\omega^b}{(\omega^2-m^2)^2+\Gamma^2}+P(\omega^2)$$
 with $P(\omega^2)=c_1+c_2\omega^2+c_3(\omega^2)^{c_4}$





comparison with other models

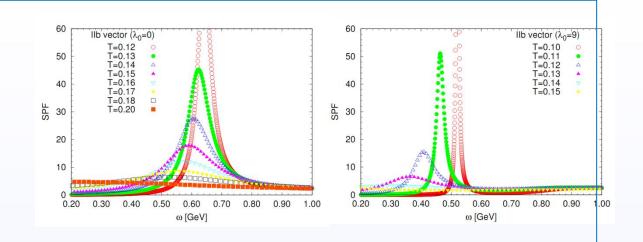
Our results:

 $\lambda = 0$:

 $T_c = 200 \,\mathrm{MeV}$

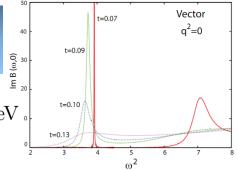
 $\lambda = 9$:

 $T_c = 150 \text{ MeV}$



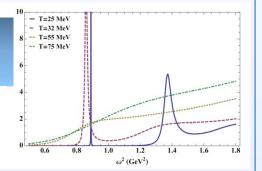


 $T_c = 0.15\sqrt{c} = 60 \text{ MeV}$



Colangelo, et al (arXiv:0909.1534)

 $T_c = 75 \text{ MeV}$





Why is this difference

 the introduction of the scale µg in the modified 5D metric will be crucial for causing the differences.

$$a^{2}(z) = 1/z^{2} + \mu_{g}^{2} \qquad \mu_{g}^{2}/(\pi^{2}T^{2})u^{2}$$

(µg couples temperature in the metric)

- dilaton and the scalar field
 - the dilaton in our model is determined as the solution of equation of motion with the given v(z) which consists of the finite temperature and the zero-temperature parts
 - to cooperate the chiral symmetry breaking and linear confinement we have to obtain the correct boundary conditions for both the VEV at the UV boundary and the dilaton at the IR boundary

